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Channel Modeling and Threshold Signal Processing in Underwater Acoustics: An Analytical Overview

David Middleton
Office of the Associate Technical
Director for Technology





Naval Underwater Systems Center Newport, Rhode Island / New London, Connecticut

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PREFACE

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W. A. VON WINKLE

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An overview of underwater acoustic channel modeling and threshold signal processing is presented which emphasizes the inhomogeneous, random, and nongaussian nature of the generalized channel, combined with appropriate weak-signal detection and estimation. Principal attention is given to the formal structuring of the scattered and ambient acoustic noise fields, as well as that of the desired signal, including both fading and doppler "smear" phenomena. The rôle of general receiving arrays is noted, as well as their impact on spatial as well as temporal signal processing and beam forming, as indicated by various performance measures in detection and estimation. The emphasis here is on limiting optimum threshold systems, with some attention to suboptimum cases. Specific first-order probability density functions (pdf's) for the nongaussian components of typical underwater acoustic noise environments are included along with their field covariances. Several examples incorporating these pdf's are given, to illustrate the (con't. on back)							
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applications and general methods involved. The fundamental rôle of the detector structure in determining the associated optimum estimators is noted: the estimators are specific linear or nonlinear functionals of the original optimum detector algorithm, depending on the criterion (i.e., minimization of the chosen error or cost function) selected. Results for both coherent and incoherent modes of reception are presented, reflecting the fact that frequently signal epoch is not known initially at the receiver.

To supplement the general discussion a selected list of references is included, to provide direct access to specific, detailed problems, techniques, and results,

for which the present paper is only a guide.

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List of Principal Symbols

 A_{Δ} = overlap index

 α , α_I , α_H , α_A = fields

 $\mathbf{Q}, \mathbf{Q}_{\mathsf{T}}, \mathbf{Q}_{\mathsf{R}} = \mathsf{beam} \mathsf{patterns}$

= (in context) a Class B noise parameter, cf. Eq. (6.13)

= input signal-to-noise ratio

= detector bias B*

= mean speed of sound

 $\hat{\eta}$ = field renormalization operator

 F_1 , F_2 , F_n , F_Q = characteristic functions

= Fourier transforms

FOM = Faure-Ol'shevskii-Middleton

FOR = Feedback operational representation

FOS = Feedback Operational Solution

 G_T = source function

G_T = source operator

 g_{J}^{\star} , $g_{J}^{}$ = threshold detector algorithms

 γ , γ_{α}^{*} = estimators

 $h, h_R, h_F = (filter)$ weighting functions

 \hat{i}_T , \hat{i}_R , \hat{i}_{oT} , \hat{i}_{oR} = unit vectors

J = MN = number of space-timefield samples

 K_x , K_α , K_T = covariance functions

K ≈ detector threshold

k = vector wave number

 $L = \hat{R} X = basic received waveform$ $\hat{L}^{(0)}$ = linear homogeneous medium

operator

L = source field

 ℓ_1 , ℓ = log-likelihood ratios

 $L^{(2)}$, $L^{(4)}$ = statistical parameters of the noise field

 \hat{M} , \hat{M}_{m} = integral homogeneous medium operators

 $\mu = p/q = ratio of a priori data$ probabilities

= number of time samples; also, number of radiating sources

dN, dn ≈ counting functionals

 $\stackrel{\vee}{\sim}$, $\stackrel{\vee}{\sim}_0$, $\stackrel{\vee}{\sim}_{OR}$ = vector spatial frequencies

ω = 2πf = angular frequency

() = operator symbol

 $P_e^{()}$, $P_D^{()}$, $P_e^{()}$, $P_D^{()}$ = error and detection probabilities

PTSS = perturbation theoretical series solutions

 \hat{Q} , $\hat{Q}^{(0)}$, $\hat{q}^{(k)}$ = linear inhomogeneous medium operators

$$\hat{R}$$
 = array operator

$$<>_{R}$$
 = radiation event average \odot = error function

$$\rho$$
 = process density

$$s = i\omega$$

$$s = i\omega$$
 $\sigma_{j}^{2}, (\sigma_{j}^{*})^{2} = \text{detector variances}$

$$\sigma$$
 = pdf or density

$$<>_S$$
 = average over space and parameters (θ)

$$T_{AT}$$
, T_{AR} , T_{M} , $T_{R_{O}}$, $T_{T_{O}}$ = transformation operators

 T_{X} , T_{OX} = relaxation times

$$\underline{\theta}$$
, $\underline{\theta}'$ = sets of parameters

$$w_1$$
 = first-order pdf

$$X(t)$$
, $x = received wave forms$

Z,
$$Z_R$$
, Z_S , Z_θ = domain of integration

Channel Modeling and Threshold Signal Processing in Underwater Acoustics: An Analytical Overview

1. Introduction

The purpose of this paper is twofold: (1) to combine a canonical characterization of random underwater acoustic fields or "channels" with the principle elements of threshold signal processing, in essence, signal detection and extraction; and (2) to provide, in the process, a necessarily concise analytic description of some of the principal formal approaches required to achieve these ends. Accordingly, our presentation is primarily tutorial, in keeping with the space available. This overview seeks to provide one possible albeit rather formal framework for handling the increasingly specialized problems of medium modeling, signal processing, and performance evaluation, which are now encountered in applications. We shall lean heavily on some of the author's recent work on channel characterization [1], [2] and space-time signal processing [3]-[5], as well as on recent related work on the statistical-physical modeling of nongaussian noise and interference fields [6]-[11]. (Tutorial background studies formally relevant to the above are provided in [12], [13] for nongaussian electromagnetic interference (EMI) environments.) In large part we shall be roncerned with optimal processing, considered as a limiting form of system structure and performance, to be approximated in practice according to the inevitable economic and operational constraints of the task at hand. [For related, suboptimum systems, we note the parallel telecommunication examples descussed in [4], [14]-[16].]

Thus, our general aim is to provide a framework for integrating the effects of the channel (or medium) and transmission and reception of signals through it, to obtain system optimization and associated performance measures. The complex acoustic channels specifically considered here are the essentially linear underwater media of typical ocean environments, which include volume and interface (surface and bottom) scatter mechanisms, and other inhomogeneities (e.g., gradients, internal waves, etc.). Ambient noise mechanisms must also be considered. These include shipping, biological phenomena, and various geophysical sources, e.g., arctic ice,

seismic activity, etc. For example, see the recent work of Dwyer [17 and refs. therein] on underice ambient noise. For the signal processing following reception we are mainly concerned here with threshold operation, up to the point of signal decoding procedures.

Because the channel exerts such a critical influence on the signals transmitted through it, it is particularly important to include the principal effects whereby these non-ideal, inhomogeneous media degrade and modify the original signal. A canonical characterization [namely, one that is invariant of any particular medium properties (until the medium is specified)] is needed in order to provide the required generality in compact and manageable form, and to allow appropriate approximations in specific cases. Similarly, it is well understood that effective signal processing must incorporate the relevant characteristics of the channel, as well as the desired signal, optimally for limiting, optimal performance and at least adequately, for practical suboptimum systems, which may be more or less close to the ideal, optimum limits.

A critical element of signal processing is spatial sampling, achieved through the array or aperture by which source and receiver are coupled to the medium. In addition to the usual temporal processing, spatial processing becomes important in reducing minimum detectable signal levels whenever the interfering noise fields are nonuniform over the array or aperture, as long as signal wavefront levels are maintained uniform [3]-[5]. This is a further reason for examining the field structure of the acoustic channel (or any chanmel, for that matter), and for including the specific effects of the coupling arrays in the overall processing program. A second critical feature is the practical fact that the relevant statistics, i.e., the various probability distributions which describe the noise fields and, in part, the desired signal fields as well, are very often highly nongaussian, characteristic of ship, underice, and biological mechanisms, among others [3], [5]. A third important concept is that of threshold signal processing, where it is possible to obtain canonical algorithms, whose forms are independent of the particular noise statistics and signal structures, and which for suitably weak signals are both locally and asymptotically (i.e., as processing gain

is increased) optimum [4], [13], [16]. Applying the specific (non)gaussian distributions indicated by these algorithms then yields the corresponding processing and performance. In practice, of course, optimal structures are limiting forms which are only approximated practically. However, using the theoretical weak-signal optimum with care to avoid destruction of the information-bearing portion of the signal, provides an effective guide to suboptimum algorithms which are "close to" the optimum and which perform effectively at all signal levels: [18], for example.

Because space here limits the detail with which we can accomplish our dual task of channel characterization and threshold signal reception for these underwater acoustic régimes, our treatment is basically a "top-down" approach, proceeding from rather general formulations to more specific examples which allow us to invoke some of the physics of the problems in question. [Extensive references assist the reader in his pursuit of analytical detail and special solutions.] One advantage of the "top-down" approach is that it readily reveals the broad connections between channel physics, effective signal processing procedures, and the general communicational tasks of signal detection and estimation in noisy environments, focused here on underwater acoustic milieux.

1.1 Organization

Accordingly, the present paper is organized as follows:

<u>Section 2</u> gives an introductory review of the general problem, including the channel and the generalized communication processes involved.

Section 3, beginning Part I: Medium Modeling, discusses some typical propagation equations for inhomogeneous deterministic and random media (Langevin equations) and their formal solutions. This includes a number of recent results developed for dealing with (linear) random media, in particular complex scattering channels like the ocean, which combines a variety of interacting scattering mechanisms. [An equivalent diagrammatic formulation is outlined in Appendix A.1.]

<u>Section 4</u> presents an outline of a second-order statistical representation of these channels, while

- <u>Section 5</u> gives two important illustrative examples for underwater acoustics: I, weak volume scatter, and II, an "exact" ocean surface scatter formalism, with a short discussion of their physical properties and structure.
- <u>Section 6</u> concludes Part I with a summary of poisson field statistics which are needed in the description of both ambient and scattered fields.
- Sections 7 and 8, introducing Part II: Acoustic Threshold Signal Processing, summarize some of the main results for threshold signal detection, including space-time sampling, performance results, and examples of optimum and suboptimum detection.
- Section 9 then provides a short summary of signal estimation procedures, which are closely related to the detection formalism in the limiting optimum cases. As in detection, these are likewise applied to threshold signal operation, including a specific example of weak signal estimation.
- Section 10 completes our analytical overview of signal processing problems associated with the underwater acoustic environment, including supporting references to the technical procedures presented in the accompanying References.

[Supplementing the above, Appendix A.1 outlines elements of diagram methods used here for the solutions of scattering problems, while Appendix A.2 provides a very concise summary of decision-theoretical formalism needed in the signal estimation aspects of reception.] Various new results, appearing in the cited references, are briefly summarized when appropriate. We stress once more the isomorphism between the channel modeling and signal processing structures of the (scalar) forms arising for telecommunications in EMI environments and that reviewed here for underwater acoustic applications. Finally, we emphasize that what is ultimately needed for an effective treatment of the general processing problem is an interdisciplinary approach, ranging from the appropriate physical models to the methods of statistical communication theory (SCT). This will be amply demonstrated in the following sections.

2. Formulation

The general situation is embodied in the operational relation [2],

[19]

$$\{v\} = I_{R_0} (I_{AR}I_M^{(N)}I_{AT})I_{T_0}\{u\}, \qquad (2.1)$$

where $\{u\}$ is a set of "messages" to be transmitted and $\{v\}$ is the ensemble of received messages, or decisions, which are consequent upon the set $\{u\}$. The J_{AT} , $J_{M}^{(N)}$, J_{AR} are general operators which describe respectively the process of coupling to the medium by the transmitting aperture, the effects of the medium itself on the resulting injected space-time signals, and the receiving aperture, by which the received field is converted into a (temporal) wave, for further signal-processing by the receiver itself. The operators J_{T_0} , J_{R_0} are temporal processes only, representing the overall "encoding" and "decoding" of the original and final "message" sets, $\{u\}$, $\{v\}$, in $\{2.1\}$. This J_{T_0} includes the actual encoding process whereby "messages" are converted to signals, and then are suitably modulated (as narrow-band waves, usually), to drive the transmitting aperture J_{AR} . Conversely, J_{R_0} includes the corresponding decoding process, as well as any appropriate (usually non-linear) signal processing [20], as shown in Fig. 2.1.

In our present study we are concerned essentially with the "signal" portions of (2.1), namely the injected signal, $S_{in}(t)$, the <u>received wave</u>, X(t), following the receiver aperture \mathcal{I}_{AR} , and ultimately here, with the detected wave $Y = \mathcal{I}_{Det}\{X\}$ (before any decoding), viz.:

$$Y(t) = T_{\text{Det}}\{X(t)\} = I_{\text{Det}}\{I_{AR}I_{M}^{(N)}I_{AT}\{S_{in}\}\} = I_{\text{Det}}I_{AR}^{\alpha}, \qquad (2.2)$$

where S_{in} is an injected (possibly encoded) signal and where

$$\alpha(\underline{R},t) = \underline{J}_{M}^{(N)}\underline{J}_{AT}\{S_{in}\} = \underline{T}_{M}^{(N)}\{-G_{T}\}$$
(2.3)

is the field at some point (\underline{R},t) in the medium in question, and G_T is a source function. Here G_T and X(t) are explicitly given by

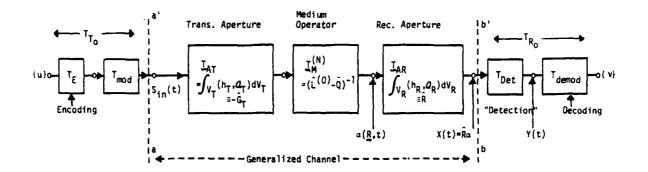


Figure 2.1 Operational schematic of a generalized channel: here, a complex, ocean medium and subsequent signal processing (T_{R_0}) .

$$G_{T} = \hat{G}_{T}S_{in} = G_{T}(\underline{\xi}, t|S_{in}) = \int_{-\infty}^{\infty} h_{T}(t-\tau, t|\underline{\xi})S_{in}(\tau,\underline{\xi})d\tau \quad (\equiv h_{T}*S_{in}),$$
(2.4a)

and

$$X(t) = \hat{R}\alpha = \int_{R} dV_{R}(\underline{R}) \int_{-\infty}^{\infty} h_{R}(t-\tau,t) \underline{R} \varepsilon V_{R}(\underline{R},\tau) d\tau = \int_{R} h_{R} \star \alpha dV_{R}, \qquad (2.4b)$$

where the source and array operators, $\hat{\textbf{G}}_{\text{T}}$, $\hat{\textbf{R}}$ are respectively \dagger

$$\hat{G}_{T} = \hat{G}_{T}(t, \underline{\xi}|\tau) \equiv \int_{-\infty}^{\infty} h_{T}(\tau, t|\underline{\xi}) \left(\underbrace{)}_{\underline{R}, t-\tau} d\tau \right; \qquad (2.5)$$

$$\hat{R} \equiv \hat{R}(t|\underline{R},\tau) = \int_{R} dV_{R} \int_{-\infty}^{\infty} h_{R}(\tau,t|\underline{R} \in V_{R}) \quad ()_{\underline{R},t-\tau} d\tau \quad (=\underline{T}_{AR}). \quad (2.6)$$

We remark that $h_{T,R}$ are respectively the filter weighting functions, now four-dimensional (i.e., spatio-temporal) quantities, of the transmitting and receiving apertures, which may be time-varying as well as frequency selective. Their space-time fourier transforms are the (complex) beam patterns

$$\mathbf{Q}_{\mathsf{T},\mathsf{R}}(\underline{\psi},\mathsf{f}|\mathsf{t}) = \underbrace{\mathbf{J}_{\xi}^{-1}}_{\mathsf{J}_{\mathsf{t}}} \mathbf{J}_{\mathsf{t}}^{\mathsf{h}} \mathbf{J}_{\mathsf{R}}^{\mathsf{R}} = \int_{\mathsf{R},\mathsf{V}_{\mathsf{T}}} \mathsf{d}\underline{\xi} e^{2\pi \mathrm{i}\underline{\psi}\cdot\underline{\xi}} \int_{-\infty}^{\infty} \mathsf{h}_{\mathsf{T},\mathsf{R}} e^{-\mathrm{i}\omega\mathsf{T}} \mathsf{d}_{\mathsf{T}}, \quad \omega = 2\pi \mathrm{f}, \quad (2.7)$$

with y a vector spatial frequency (which may itself be a complex function of frequency, f), when the medium is absorptive.

 $[\]dagger$ Throughout, we denote operators by (^).

2.1 The Generalized Channel

For the linear media assumed throughout, the field α obeys a partial integro-differential equation of the form [2]

$$(\hat{L}^{(0)} - \hat{Q})\alpha = -G_T + [b.c.'s + i.c.'s],$$
 (2.8)

where $\hat{L}^{(0)}$ is a linear (scalar) partial differential operator with $\hat{L}^{(0)}$ associated with the homogeneous portion of the medium. Here Q is, in general, a (scalar, linear) integro-differential scattering operator, which describes the interaction of the incident (homogeneous) field with the (differential, or local) scattering (i.e., re-radiating elements) of the inhomogeneous portion of the medium. Boundary (b.c.) and initial conditions (i.c.) are, of course, a necessary part of (2.8), as indicated.

Since we are dealing generally with <u>random media</u>, inasmuch as $\hat{\mathbb{Q}}$ contains random as well as deterministic components, we are concerned with the appropriate <u>ensembles</u>, $\{\alpha\}$, $\{X\}$, etc. The ensemble of equations (2.8) governing propagation of the field is now the <u>Langevin equation</u> ([2]; [20], chapter 10). Since, formally,

$$\{\alpha = (\hat{L}^{(0)} - \hat{Q})^{-1} (-G_T)\},$$
 (2.9)

comparison with (2.3) shows that

$$\{T_{M}^{(N)} = (\hat{L}^{(0)} - \hat{Q})^{-1}\}. \tag{2.10}$$

This, and Eqs. (2.2), (2.3), are shown schematically in Fig. 2.1.

For the purely deterministic special cases we seek solutions of (2.8) directly. However, for the general situation of random media, the "solutions" of the corresponding Langevin equation (2.9) are the various statistics of the field $\{\alpha\}$, e.g., the various moments $<\infty$, $<\alpha_1\alpha_2>$, etc., and the moments of the (linear functionals) of the field, $\{X(t)\}$, cf. (2.4b), e.g., <X>, $<X_1X_2>$, etc., and more comprehensively, the various probability densities (and distributions) of α and X. Thus we are interested here primarily in the means $<\infty$, <X>, the intensities $<\alpha_1^2>$, $<X_2^2>$, and the covariances ($<\alpha_1^2\alpha_2>$, $<<X_1^2A_2>$). [Appendix A.1 provides a short introduction to various methods of evaluating $<\infty$, $<\alpha_1\alpha_2>$, etc.] Accordingly,

for the purposes of this paper, we shall define the <u>generalized channel</u> here as the sequence of operations indicated by (2.2), viz. X(t), or in a less restricted sense, by the resultant field α , (2.3), cf. Fig. 2.1. Consequently, the desired statistical description of the generalized channel is just that of $X = \hat{R}\alpha$, or of α , namely, the various above-mentioned moments.

When X is a gaussian process, these first- and second-order moments are sufficient for subsequent signal processing (T_{R_0} in Fig. 2.1), to yield {v}. But, as is often the case, X here is <u>not</u> gaussian, so that at least first-order <u>distributions</u> of X must be developed for effective processing. This is discussed further in Section 6.2 below.

2.2 The Rôle of Scattering Theory

The stochastic character of the medium or channel (as above), arises, of course, from the random spatio-temporal inhomogeneities within and bounding the medium. Our quantitative description of the medium now as a generalized communication channel accordingly must be made in terms of an appropriate scattering theory, particularly one which is capable of including both the geometrical and statistical features of real-world situations in manageable fashion. We include formally as a special case of scattering models that of ambient noise, where now the secondary radiating or scattering sources are replaced by primary sources, cf. Sec. 6.1.

Some recent new approaches [1] provide additional technical and insightful methods for "solving" the Langevin equations (2.9) of this general underwater acoustic channel. These include:

- I. The concept of <u>Feedback Operational Solutions (FOS)</u>, cf. Sec. 3.2 ff., to represent the canonical Langevin equation (2.9);
- II. The ability to handle possibly <u>interacting scattering modes</u> (the so-called "M-form") (Sec. 2.5 of [1]); and
- III. The observation that <u>boundaries</u>, wave surface and ocean bottoms for example, <u>can be treated as distributed inhomogeneities</u> in an otherwise infinite medium (Sec. 2.2, [1]).

The first item (I) relates the analytic solutions and diagram methods to a generalized, i.e., four-dimensional, feedback theory and formalism; the second (II) offers a simple taxonomy for accounting for any significant

scattering interaction, e.g., surface with volume scatter, etc.; and the third (III) greatly simplifies the analysis, at least in the usual farfield (or "radiation-field") cases, by effectively converting distributed (or integral) boundary conditions into local, or differential boundary conditions, i.e., plane-wave, at-a-point b.c.'s.

Finally, we must note that scattering theory, particularly with respect to acoustic and electromagnetic propagation, is a venerable field of study, with a correspondingly vast literature. A full citation of the principal publications to date is well beyond our capabilities here. However, a useful and extensive summary of modern "classical" methods and results, chiefly in the electromagnetic cases, has recently been given by Ishimaru [22], along with extensive references. An earlier and more mathematical discussion of scattering problems has been provided in the interesting review paper by Frisch [23]. For acoustical scattering, particularly in the ocean and at its interfaces, the somewhat later review articles by Fortuin [25] and Horton [26] regarding surface scatter are especially noted. In the latter connection, see also [11], [43]-[56]; in particular, Bass and Fuks [48] and [43]-[45], [47], [49], and [50], for wave-surface scattering. Other important recent studies, devoted principally to volume scatter, are described by Flatté et al. [46]. Tatarskii's significant work [24] is also well known and is also especially appropriate to the treatment of both acoustic and electromagnetic scattering in the atmosphere, and contains, as well, many useful references.

An important departure from "classical continuum theory" [25], [26] is the so-called FOM theory (after Faure [28], Ol'shevskii [29], and Middleton [31]. This latter is based on first-order, i.e., independent poisson point-scatter models. Various elements of the author's more recent approaches [2], which include both the "classical" and FOM theories as particular components of the general multiple-scatter régime, cf. Appendixes A.1, A.2 have already been presented in different stages of evolution and development since 1973, [32], [40]. The essentially mature form is described in [2].

Part I. Medium Modeling

Inhomogeneous Stochastic Media: Langevin Equations and Solutions

Starting with the appropriate equations of state, mass, and momentum conservation, one can obtain by suitable approximations ([2], [21]) of the resulting nonlinear propagation equations, linearized "wave equations," which are special variants of (2.8). It is more convenient, however, to proceed with the canonical form (2.8) directly. Here $\hat{L}^{(0)}$ is a linear differential operator, with cöefficients independent of (R,t), while \hat{Q} , also linear, may be an integro-differential (or global) operator and is dependent on position (R,t). (This character of \hat{Q} arises from scattering element doppler, cf. [36].) For the moment we regard \hat{Q} , and hence (2.8), as deterministic.

An important example for acoustic propagation in underwater media is the case of inhomogeneous, absorptive media [2], for which (2.8) is specifically [33], [24]

$$\{(1 + \tau_{\mathsf{X}}(\underline{\mathsf{R}},\mathsf{t}) \, \frac{\partial}{\partial \mathsf{t}}) \nabla^2 - \frac{1}{\mathsf{c}_0^2} \, [1 + \varepsilon(\underline{\mathsf{R}},\mathsf{t})] \, \frac{\partial^2}{\partial \mathsf{t}^2} \}_{\alpha} = -\mathsf{G}_{\mathsf{T}}, \tag{3.1}$$

where

$$\therefore \hat{L}^{(0)} = (1 + \tau_{ox} \frac{\partial}{\partial t}) \nabla^2 - \frac{1}{c_o^2} \frac{\partial^2}{\partial t^2}; \quad \hat{Q} = \frac{\varepsilon}{c_o^2} \frac{\partial^2}{\partial t^2} - \tau_{ox} \gamma_x \frac{\partial}{\partial t} \nabla^2,$$
(3.1a)

with $\tau_{X} \equiv \tau_{OX} \cdot (1 + \gamma_{X}[\underline{R}, t])$ generally. When $\tau_{OX} = 0$, (3.1) reduces to the more familiar "extended" Helmholtz equation

$$\{\nabla^2 - \frac{1}{c_0^2} \left[1 + \varepsilon(\mathbf{R}, \mathbf{t})\right] \frac{\partial^2}{\partial t^2}\}\alpha = -G_T, \tag{3.2},$$

with

$$\hat{Q} = \frac{\varepsilon}{c_0^2} \frac{\partial^2}{\partial t^2} ; \qquad \hat{L}^{(0)} \equiv \nabla^2 \frac{1}{c_0^2} \frac{\partial^2}{\partial t^2} . \qquad (3.2a)$$

Here

$$\underline{R} = \hat{\mathbf{I}}_{X} \times + \hat{\mathbf{I}}_{Y} + \hat{\mathbf{I}}_{Z} z \tag{3.3}$$

and rectangular cöordinates are assumed throughout, so that $\nabla \cdot \nabla = \nabla^2$, etc. Specifically, $c_0 = (\text{constant})$ wave-front speed, ε embodies the effects of velocity gradients, internal wave phenomena, and/or (weak) local turbulence; while τ_X represents the effects of relaxation absorption (due to Mg₂SO₄ and other salts).

With inhomogeneous (but still deterministic) media conventional techniques of solution fail, largely because the medium is not reciprocal, and there is no general way of applying and evaluating the initial conditions over the various volume integrals which appear in the development of the generalized Huygen's principle (GHP) when $\hat{Q} \neq 0$ (\bar{L}^2). Moreover, since these media are space-time variable in their inhomogeneities, e.g., $\hat{Q} = \hat{Q}(R,t|\ldots)$, either locally, or because of local doppler [36], such media do not support space- and time-harmonic solutions [2], nor are the standard perturbational and variational techniques of the "classical" approaches [35] generally applicable, particularly when $\hat{Q} = \hat{Q}(R,t|\ldots)$ and is a random operator, as will be the case here. Finally, in addition to all these difficulties is the usually insurperable one of bounded media when the boundaries themselves are random and moving.

3.1 Random Media

Accordingly, entirely new approaches are required, for manageable solutions of (2.8) generally (and (3.1) in particular), which embody the physical realities of the application in question ([2], [21]), and which, in particular, include both the random character of \hat{Q} , i.e., $\epsilon(R,t)$ in (3.1) etc., and the random doppler effects [36], as well. These involve incorporating the effects of boundaries into the now random inhomogeneity operator \hat{Q} , as well as different types of scatter mechanisms, e.g., $\hat{Q} = \hat{Q}_S + \hat{Q}_{vol} + \hat{Q}_{Bottom}$, etc., with $\hat{Q}_{vol} = \hat{Q}_{Inh} + \hat{Q}_{Disc}$, where \hat{Q}_{Inh} represents the continuously (random) variations in medium properties, while \hat{Q}_{Disc}) embodies the effects of localized, or discrete particles, bubbles, etc. A typical geometry is sketched in Fig. 3.1, showing the usual variety of interactions.

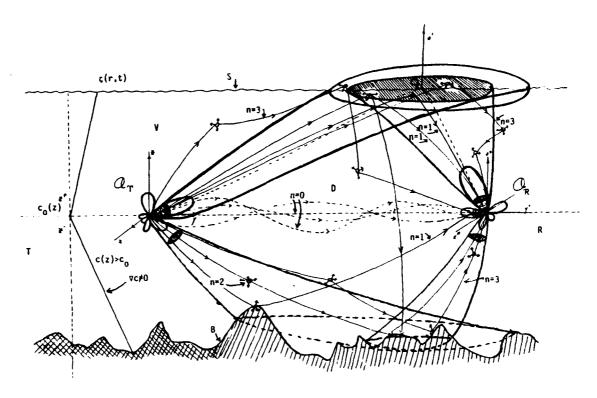


Figure 3.1: Schematic scatter channel geometry of source (T), medium (V), and boundaries (S,B), with receiver (R), D = unscattered, or "direct" path, and n is the order of the interactions. Sufficiently high frequencies are assumed to permit illustrative ray paths.

For these linear media, with $\hat{L}^{(0)}^{-1} = \hat{M}$ (which also includes boundary and initial conditions), it is easily seen that, in operator form, (2.8) becomes, for each member of the ensemble of propagation equations:

$$(\hat{1}-\hat{M}\hat{Q})\alpha = \hat{M}(-G_{\mathsf{T}}) ; \quad \alpha = \alpha_{\mathsf{H}} + \alpha_{\mathsf{I}} ; \quad \alpha_{\mathsf{H}} = \hat{M}(-G_{\mathsf{T}}) , \qquad (3.4)$$

where $\hat{1}_{\alpha}$ = α , etc. We note that α_H , (3.4), is the unscattered field, or field of the purely deterministic, i.e., homogeneous portion of the medium here. Moreover, because any boundaries are included in \hat{Q} , \hat{M} is replaced by \hat{M}_m , so that now

$$\alpha_{\rm H} = \hat{M}_{\infty}(-G_{\rm T}) , \qquad (3.4a)$$

where the infinite domain operator for the homogeneous component of the medium becomes

$$\hat{\mathbf{M}}_{\infty} = \hat{\mathbf{M}}_{\infty}(\underline{\mathbf{R}}, \mathbf{t}|\underline{\mathbf{R}}', \mathbf{t}') = -\int_{-\infty}^{\infty} d\mathbf{t}' \int_{(\infty)} \underline{\mathbf{d}}\underline{\mathbf{R}}' \, \mathbf{q}(\underline{\mathbf{R}}, \mathbf{t}|\underline{\mathbf{R}}', \mathbf{t}')_{\infty}()_{\underline{\mathbf{R}}', \mathbf{t}'} (3.5a)$$

$$= -\mathbf{3}_{S}^{-1}(\hat{\mathbf{Y}}_{0,\infty}) = -\mathbf{3}_{K}\mathbf{3}_{S}^{-1}\{\hat{\mathbf{y}}_{0,\infty}\}. \tag{3.5b}$$

Here

$$\hat{\mathbf{y}}_{0,\infty} = \hat{\mathbf{y}}_{0}(\underline{\mathbf{k}},s) \int dt' \int_{(\infty)} e^{-i\underline{\mathbf{k}}\cdot\underline{\mathbf{R}}'-st'} ()_{\underline{\mathbf{R}}',t'} d\underline{\mathbf{R}}',$$

$$\hat{\mathbf{y}}_{0,\infty} = \hat{\mathbf{y}}_{0,\infty}(s|\rho) \int dt' e^{-st'} ()_{t'}$$
(3.5c)

and

$$\underline{k} = \hat{1}_{x} k_{x} + \hat{1}_{y} k_{y} + \hat{1}_{z} k_{z} = 2\pi y$$
 (3.5d)

is a vector wave number, with $\underline{\vee}$ a vector spatial frequency, cf. (2.7). Here g_{∞} is the green's function of the corresponding infinite, homogeneous medium ($\hat{Q} = 0$). For the ocean medium supporting propagation according to (3.1a), we have specifically the (homogeneous) operator (kernels)

$$Y_{0,\infty} = \frac{e^{-\rho \frac{S}{C_0} (1+\tau_{0X}S)^{-\frac{1}{2}} - St'}}{4\pi\rho (1+\tau_{0X}S)^{\frac{1}{2}}}; \qquad (3.6a)$$

$$y_{0,\infty} = [k^2(1+\tau_{0x}s) + s^2/c_0^2]^{-1}; \quad \rho = [R-R']; \quad k^2 = \underline{k} \cdot \underline{k}. \quad (3.6b)$$

For the extended Helmholtz equation (3.2), (3.2a), Eqs. (3.6) for the homogeneous part reduce to the simpler relations

$$Y_{0,\infty} = \frac{e^{-\rho s/c_0 - st'}}{4\pi\rho}; \quad y_{0,\infty} = (k^2 + \frac{s^2}{c_0^2})^{-1};$$

$$\therefore M(\underline{R}, t|\underline{R}', t')_{\infty} = \frac{\delta(t - t' - \rho/c_0)}{4\pi\rho}.$$
(3.7)

3.2 Field Series and Feedback Solutions

The ensemble solution of (3.4) is now given formally by [2], [21]):

$$\alpha = (\frac{\hat{1}}{\hat{1} - \hat{M}_{\infty} \hat{Q}}) \alpha_{H} = (\hat{1} - \hat{\eta}_{\infty})^{-1} \alpha_{H} ; \qquad \hat{\eta} = \hat{M}_{\infty} \hat{Q} , \qquad (3.8)$$

where $\hat{\mathbf{\eta}}$ is now the random <u>field renormalization operator (FRO)</u>, when considered over all member equations of the Langvin equation. As long as $\|\hat{\boldsymbol{\gamma}}\| < \infty$ (for each member equation), (3.8) supports a series expansion:

$$\alpha = \alpha_{\mathsf{H}} + (\frac{\hat{\eta}}{\hat{1} - \hat{\eta}}) \alpha_{\mathsf{H}} = \alpha_{\mathsf{H}} + \sum_{\mathsf{n}=1}^{\infty} \hat{\eta}_{\infty}^{(\mathsf{n})} \alpha_{\mathsf{H}} : \|\hat{\eta}\| < 1;$$
 (3.9)

(or
$$\alpha = -\sum_{n=0}^{\infty} (\hat{\eta}^{-1})^{(n+1)} \alpha_{H}, \quad 1 < ||\hat{\eta}|| < \infty$$
). (3.9a)

(Usually it can be shown that $\|\hat{\boldsymbol{\gamma}}\| < 1$, [2]; we remark, also, that $\hat{\mathbb{M}}_{\infty}$ and $\hat{\mathbb{Q}}$ do not usually commute, but commutation can always be achieved in the proper transform space, usually ($\underline{\mathbf{k}}$,s)-space [2].) The expansion (3.9), or (3.9a), is called a <u>Perturbation Theoretical Series Solution</u> (PTSS), with the convergence condition as indicated formally above.

Equations (3.8), (3.9) are exact (ensemble) solutions. Eq. (3.8), moreover, is called the <u>Feedback Operational Solution</u> (FOS), since $\hat{\mathbb{M}}_{\infty}$ may be interpreted as a "feedforward" and $\hat{\mathbb{Q}}$, a "feedback" operator, cf. [2], [11], in the manner of Fig. 3.2. In fact, the <u>PTSS</u> in (3.9), expressed as a series of loop iterations formally, indicates computationally how exact numerical solutions may be approached. The evaluation of the <u>PTSS</u> may then be viewed as both a problem in modern control and computational theory, extended now to four dimensions and governed by partial (rather than ordinary) differential equations and influenced by a radiation condition (on $\hat{\mathbb{M}}_{\infty}$), as well as boundary conditions applied locally to $\hat{\mathbb{Q}}$.

However, the formal solutions (3.8), (3.9) are not very useful in their present form since $\hat{\bf \Omega}$ is still a random operator. To achieve useful results the Langevin equations must be converted into various

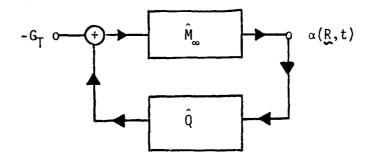


Figure 3.2 Feedback Operational Representation (FOR) for the propagation Eq. (3.8), for linear inhomogeneous media.

deterministic forms, which represent the various <u>moments</u> of the scattered field, as we have noted above in Sec. 2.1. This is done in the manner described in Appendix A.1 and leads to suitable replacement of the random \hat{Q} by an appropriate deterministic operator $\hat{Q}_1^{(d)}$, $\hat{Q}_{12}^{(d)}$, etc. for the first-, second-, and higher-order moments of the field α , cf. Secs. A.1-1. Thus, we get a <u>Dyson equation</u> for α , a form of the Bethe-Salpeter equation for $\alpha_1 \alpha_2 >$, etc. For example, the FOR of Fig. 3.2 applies now, with α replaced by α and α by α and α by α and α by α approximating α the various terms of the PTSS can be evaluated, although this is not an easy matter.

Finally, we observe that the ensemble solutions (3.8), (3.9) may also be insightfully described by equivalent <u>diagram representations</u>, which are also particularly useful in suggesting appropriate, simplifying approximations. A concise summary of this approach is provided in Appendix A.1, some results of which we use below in the text, cf. Sec. 5 ff. Related methods, involving path integrals and various models of distributed ocean inhomogeneities, with particular attention to long distance sound transmission in the ocean, are treated in the recent important studies of Flatté et al. [46]. These offer a variety of techniques for realizing the inhomogeneity operator \hat{Q} and its associated statistics (see, for example, Chapter 10 of [40].)

4. Channel Characterization and Statistics

The (random) channel operator $\mathcal{I}_{M}^{(N)}$ (= $(\hat{1}-\hat{\eta})^{-1}\hat{M}$), cf. (3.8), is next conveniently written

$$\begin{bmatrix}
I_{M}^{(N)} = \left\{\hat{1} + \frac{\hat{\eta}}{\hat{1} - \hat{\eta}}\right\} \hat{M} = \hat{M} + \hat{I} = \hat{\eta}
\end{bmatrix}; \quad \hat{I} = \frac{\hat{\eta}}{\hat{1} - \hat{\eta}} \hat{M}$$
(4.1)

where we separate the homogeneous $(\hat{M} = \hat{M}_{\infty} \text{ here})$ from the scatter operator, \hat{I} , cf. (3.4) et seq. We note that for <u>ambient fields</u>, in addition to the desired signal source (S_{in}) , we replace the source function $-G_T$, in the Langevin equations (2.8), (3.4), by $-G_T - G_A$, where now G_A is a localized, or a distributed, ambient "noise" source with an associated field, $\alpha_A = \hat{M}_{\infty}(-G_A)$. In any case, it is at once evident that the medium operators \hat{M} , \hat{I} ($\pi \eta$) are unchanged.

4.1 Homents of the Received Field

Mostly for reasons of analytical complexity, channel characterization is usually formulated in terms of the lower moments, e.g., the mean and covariance operators associated with \hat{I} . The pertinent moment operators here are, directly from (4.1),

$$\hat{\mathbf{M}} = \hat{\mathbf{M}} + \langle \hat{\mathbf{I}} \rangle , \quad \hat{\mathbf{K}}_{\mathbf{I}} = \langle \hat{\mathbf{I}}_{1} \hat{\mathbf{I}}_{2} \rangle - \langle \hat{\mathbf{I}}_{1} \rangle \langle \hat{\mathbf{I}}_{2} \rangle ;$$

$$\hat{\mathbf{M}}_{1} \hat{\mathbf{M}}_{2} = \hat{\mathbf{M}}_{1} \hat{\mathbf{M}}_{2} + \hat{\mathbf{M}}_{1} \langle \hat{\mathbf{I}}_{2} \rangle + \hat{\mathbf{M}}_{2} \langle \hat{\mathbf{I}}_{1} \rangle + \langle \hat{\mathbf{I}}_{1} \hat{\mathbf{I}}_{2} \rangle ,$$

$$(4.2)$$

and with higher moments similarly determined. These moment operators are, of course, deterministic, as a result of averaging. (Note that $\hat{I}_{1,2} = \hat{I}(\underline{R}_{1,2},t_{1,2}|\underline{R}',t')$, etc., (\underline{R}_1,t_1) , (\underline{R}_2t_2) being (generally) different points in the field.)

Expressions for the corresponding moments of the received wave X(t) are found from

$$X(t) = \hat{R}(t) = \hat{R}(\hat{M} + \hat{I})(-G_T - G_A) = \hat{R}\left[\hat{I} + \frac{\hat{\eta}}{1 - \hat{\eta}}\right](c_H + c_A).$$
 (4.3)

Thus, the mean received waveform and various second moments are respectively

$$\langle X \rangle = \hat{R} \langle \alpha \rangle = \hat{R} (\hat{M}_{\infty} + \langle \hat{I} \rangle) (-G_T - G_A)$$
 (4.4)

$$\langle x_1 x_2 \rangle = \hat{R}_1 \hat{R}_2 \langle \alpha_1 \alpha_2 \rangle = \hat{R}_1 \hat{R}_2 \hat{m}_1 \hat{m}_2 G_1 G_2, \quad G = G_T + G_A,$$
 (4.5)

with

$$\begin{array}{l} \kappa_{x}(t_{1}t_{2}) = \langle x_{1}x_{2}\rangle - \langle x_{1}\rangle \langle x_{2}\rangle = \hat{R}_{1}\hat{R}_{2}(\langle \alpha_{1}\alpha_{2}\rangle - \langle \alpha_{1}\rangle \langle \alpha_{2}\rangle) \\ = \hat{R}_{1}\hat{R}_{2}\hat{\kappa}_{1}(G_{1}G_{2}), \end{array} \tag{4.6}$$

where from (3.8) and (3.9), and (4.1), we see that <u>all orders of moments</u> of the inhomogeneity operator \hat{Q} appear in the above.

Equivalent diagram representations for the scatter operator \hat{I} and field covariance operator \hat{K}_{I} , (4.2), are given in Appendix A.1.

4.2 Remarks on Arrays

The array operators (2.4b) and (2.6), which appear in (4.3) et seq. for the received wave, can be rewritten alternatively as

$$\hat{R} \equiv \int_{-\infty}^{\infty} d\tau \int_{R} ()h(\underline{n}, t-\tau)_{R} \underline{dn}, \quad dV_{R} + \underline{dn}. \qquad (4.7)$$

Here V_R once more is the region occupied by the physical array itself (in the field α) and n locates the array element dn with respect to 0_R , the coördinate origin of (here) the receiving array, cf. Fig. 4.1. The array is a space-time linear filter.

In practice, most physical arrays consist of an assemblage of $m=1,2,\ldots,M$ discrete, but distributed elements, each of which samples the radiation field, α . Thus, the array operator (4.7) may be expressed in more detail formally as a vector operator

$$\hat{R} = \int_{-\infty}^{\infty} d\tau \int_{R_m} ()h_m(\underline{n}, t-\tau)_R d\underline{n} = {\hat{R}_m}, \quad \underline{n} \in {R_m}, \quad (4.8)$$

which defines each "component," \hat{R}_m . The relation (4.8) may be simplified further to the idealized but still very useful situation of "point" sensors, e.g., $L_{max}^{(m)} << \lambda_o$: maximum sensor dimension (each m) is always much less than the acoustic wavelength. Thus, (4.8) becomes

$$\hat{R}_{\text{"point"}} = \{\delta(\underline{\eta} - \underline{\eta}_{\text{m}}) \int_{-\infty}^{\infty} d\tau \, h_{\text{m}}(t - \tau)_{\text{R}}\}, = h_{\text{mn}-\text{R}}\delta(\underline{\eta} - \underline{\eta}_{\text{m}})\}, \quad (4.9)$$

all m; where

$$h_{mn-R} \equiv \int_{-\infty}^{\infty} ()h_{m}(\underline{\eta}_{m}, t_{n}^{-\tau})_{R}^{d\tau}$$
(4.9a)

for temporal sampling $t_n = n\Delta t$; n = 1, ..., N, as well. It is convenient to use j = m, n = 1, ..., J = MN as a single index for space-time sampling, so that now

$$\hat{R} = \{\hat{R}_{j=m,n}\} = \{\hat{R}_{j}\}$$
 (4.10)

is a J-component vector. Consequently, the <u>sampled</u>, received field X in (4.3) becomes explicitly

$$\underbrace{X = \{X_j\} = \hat{R}\alpha = \{\int_{-\infty}^{\infty} h(\underline{n}_m, t_{n}^{-\tau})_R \alpha(\underline{R}_m, \tau) d\tau\}}_{=\infty}; \quad \underline{R}_m = \underline{R}_0 + \underline{n}_m.$$
(4.11)

Thus, $X = \{X_{j=m,n}\}$ is the received data vector, after linear space-time filtering and sampling. Equation (4.11) is the basic relation, given the field α , which constitutes the received, sampled waveform. This, in turn, is the fundamental input to the "signal-processing" systems noted in Fig. 2.1, which are the principal topic of Part II following.

For narrow-band signals in the medium we can provide a more detailed structure to $h_m(\)_R$ above. In fact, by combining, i.e., summing, the m (=1, ..., M) outputs of each receiving sensor with appropriate delays $\Delta \tau_m$, we can form a beam (in the far field of the incoming signal), e.g.,

$$\underline{X} = \{X_n\} = \{\sum_{m} X_j\} = \{\sum_{m} A_R^{(m)}(\underbrace{v_0} - \underbrace{v_{oR}}; f_0)\}_{\ell_m} (t_n - \ell_{f_m})\}, \qquad (4.12)$$

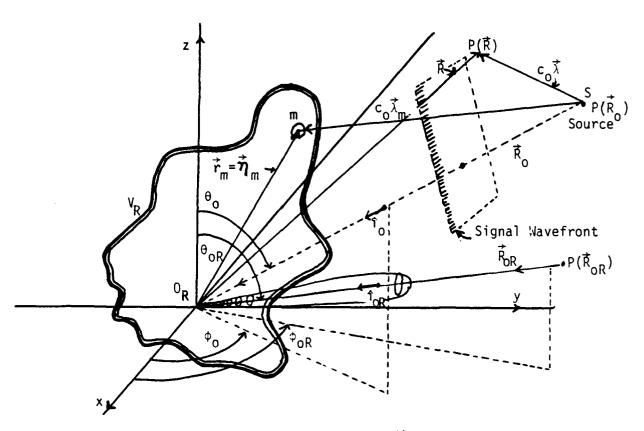


Figure 4.1 Geometry of array (in V_R), and $m^{\frac{th}{-}}$ - sensor, with beam in direction $-\hat{i}_{OR}$.

where

$$A_{R}^{(m)} = |A_{R}^{(m)}(f_{o})|e^{2\pi i (v_{o} - v_{o}_{o}_{R}) \cdot r_{m} - i\phi_{R}^{(m)}(f_{o})};$$

$$\Delta \tau_{m} = \tau_{m} - \hat{\tau}_{m} = (\hat{\tau}_{o} - \hat{\tau}_{o}_{R}) \cdot r_{m}/c_{o},$$

$$(4.12a)$$

in which $\nu_{oR} = \hat{\mathbf{1}}_{oR} f_o/c_o$, $\nu_o = \hat{\mathbf{1}}_o f_o/c_o$ are wave numbers, respectively associated with the beam direction $\hat{\mathbf{1}}_{oR}$ ($\equiv \frac{R}{oR}/|\frac{R}{oR}|$) and the direction of the field's wave-front, $\hat{\mathbf{1}}_o = \frac{R}{o}/|\frac{R}{o}|$. The geometry of the mth-sensor, the beam, and signal wavefront are shown in Fig. 4.1, where specifically the unit vectors $\hat{\mathbf{1}}_o$, $\hat{\mathbf{1}}_{oR}$ are given by

$$\hat{\mathbf{y}}_{o}(R) = \hat{\mathbf{y}}_{cos\phi}(R)^{cos\phi} o(R) + \hat{\mathbf{y}}_{sin\phi}(R)^{cos\phi} o(R) + \hat{\mathbf{z}}_{zsin\phi}(R).$$
(4.13)

In (4.12), $|A_R^{(m)}|$ and $\phi_R^{(m)}$ are the amplitude and phase of the $m\frac{th}{c}$ -sensor's response, at the center frequency (f_0) of the received signal, here the field α . For a <u>vertical array</u> of $2M_0+1=M$ sensors, for example, we have

$$\underline{r}_{m} = \underline{\hat{1}}_{z}^{m \triangle \ell}, \qquad (-M_{o}, \dots, m, \dots, M_{o})$$

$$\therefore \Delta \tau_{m} = \frac{m \triangle \ell}{c_{o}} (\sin \theta_{o} - \sin \theta_{oR}), \qquad (4.14a)$$

in which $\Delta \ell$ is the sensor spacing. For a similar, but <u>horizontal array</u>,

$$\underline{r}_{m} = \hat{\mathbf{j}}_{x}^{m} \Delta \ell \text{ and } : \Delta \tau_{m} = \frac{m \Delta \ell}{c_{o}} [\cos \phi_{o} \cos \theta_{o} - \cos \phi_{o} \cos \theta_{o} R].$$
 (4.14b)

For a narrow-band signal field $\hat{1}_0$ is usually known, or findable. Thus, the "beam" can be steered, that is, pointed in the direction of the desired signal, by choosing $\hat{1}_{OR} = \hat{1}_0$, or setting the path delays τ_m in Fig. 4.1 such that $\Delta \tau_m = 0$, cf. (4.12a). However, for an ambient or scattered noise field, each of which is the superposition of a number of possible (undesired) signal fields randomly phased with respect to one another, $\hat{1}_0$ is randomly distributed in space. There is consequently no distinct noise field wavefront or wavefront direction, and $\hat{1}_0$ must be averaged over in suitable fashion. [See Secs. 8.2 and Sec. 9 of [4], for example; also, (6.10b) and Sec. 7.3 ff.]. This randomness of $\hat{1}_0$ in the case of noise fields usually leads to spatial non-uniformity of the interfering field, which can often be exploited to increase the processing gain in signal reception, by effectively increasing the number of independent spatial noise samples vis-à-vis the (nearly) uniform desired signal field, cf. Sec. 10 of [4].

5. Ocean Volume and Surface Channels:

The Received Field X

Both the important cases of weak volume scatter in the ocean and ocean surface scatter are well approximated by at most the scattering element interacting with the homogeneous field, $\alpha_{\rm H}$, the former because multiple scatter effects are usually quite ignorable, the latter from the physical geometry of the wave surface, which likewise discourages

multiple scatter except within a typical wave crest-to-crest domain, cf. Fig. 3.1, (n = 1). (Similar considerations usually apply for ocean bottoms, as well, cf. Fig. 3.1. We shall, however, consider only the modes M = S,V, here), observing also that O(30-40 db) they are effectively independent ([21], thesis): there is negligible coupling in the ocean between surface (bottom) and volume scatter.)

Accordingly, a 1st-Born approximation suffices (as indicated by the dotted lines in (A.1d), (A.2a), and we write for the received field, X, from (A.1d):

$$X = \hat{R}\alpha = \hat{R} \longleftrightarrow \hat{T} + \xrightarrow{\hat{Q}_{M}} \hat{T}) = \hat{R}(\alpha_{H} + \hat{M}\hat{Q}_{H}\alpha_{H}), \qquad (5.1)$$

$$\longrightarrow : \hat{M} \longleftrightarrow : \hat{Q}_{M} \qquad M = S, V.$$

$$\hat{T} = -G_{T}$$
The mean wave, $\langle X \rangle$, is, with $\langle \hat{Q}_{M} \rangle = \Phi$, cf. (A.2a):

The second-moment of X, (5.1) in (4.5), is found to be, on averaging over the phases (ϕ) of the input signal.

where

$$\begin{array}{l}
0 = \hat{q} = \hat{Q} - \langle \hat{Q} \rangle_{R} = \longleftarrow 0 - 0 = \mathbf{X} \{ v^{(0)} = \sum\limits_{k=1}^{S} (\langle v^{(k)} \rangle_{R} + \Delta v^{(k)}) \}; \\
\hat{Q}^{(0)} = \langle \hat{Q} \rangle_{R} = 0, \quad \therefore \langle \hat{Q}^{(0)} \rangle_{S} = \langle \hat{Q} \rangle = \bullet = \langle 0 \rangle_{R}. \\
\hat{Q}^{(0)} = \langle 0 \rangle_{S: \text{space av.}} : \quad \mathbf{Q}^{(0)} = \langle 0 \rangle_{S} = \langle \hat{q}_{1} \hat{q}_{2} \rangle \text{ etc.} \quad \mathbf{Q}^{(0)} = \langle 0 \rangle_{R}. \\
\hat{Q}^{(0)} = \langle 0 \rangle_{S: \text{space av.}} : \quad \mathbf{Q}^{(0)} = \langle 0 \rangle_{R}. \\
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\hat{Q$$

via (A.2-6). Note that the (total) average, <>, as in Appendix A.1 et seq., involves both an average over radiation events and an average over random spatial (S) and parameter values ($\frac{1}{10}$), embodied in $<>_{S}$, cf. (5.4) and, particularly, Appendix A.2.

It is the Decomposition Principle, embodied in Q_M , (M=S,V,B), Eq. (A.2-5) et seq., which allows us to resolve the inhomogeneity operator into sets of distinct and (statistically) independent entities as specifically exhibited in Eqs. (5.3). Moreover, and most important, the entities $(k=0,\ k=1,\ k\geqslant 2)$ in (5.3) above (and in (5.1), with the help of (5.4)) have an explicit physical interpretation: Thus, the terms k=0 contain both the <u>coherent</u> and incoherent radiation contributions, from all orders $(k\geqslant 1)$ of radiation interactions, e.g., single-and multiple scattering $(k\geqslant 2)$, for example. Consequently, if $v^{(k)}$ represents the density of (illuminated) k-coupled scattering elements $(k\geqslant 1)$, then $< v^{(k)} >_R$ is the coherent radiative contribution, where $< >_R$ is the average over the radiation events associated with the ensemble of potential (re-) radiating sources. In a similar way

$$\Delta v^{(k)} \equiv v^{(k)} - \langle v^{(k)} \rangle_{R}$$
 (5.5a)

is the <u>fluctuation</u> in the density of k-coupled radiating elements, and is always associated with incoherent radiation. Accordingly, we have

$$v^{(0)} = \sum_{k=1}^{\infty} (\langle v^{(k)} \rangle_{R} + \Delta v^{(k)}) = \sum_{k=1}^{\infty} v^{(k)} = v, \qquad (5.5b)$$

$$dN^{(0)} = \sum_{k=1}^{\infty} (\langle dN^{(k)} \rangle_{R} + dn^{(k)}) = \sum_{k=1}^{\infty} dN^{(k)}.$$
 (5.5c)

The relations (5.6) et seq. obviously apply if one replaces ν etc. by dN; see also Sec. 6.1 following.

Since the medium in question is linear, the scattering or inhomogeneity operator, \hat{Q} , is a linear functional of the $v^{(k)}$, i.e., of $v^{(0)}$, e.g., $\hat{Q} = \boldsymbol{\mathcal{X}}(v^{(0)})$, thus containing all types of radiation interaction effects. Moments of $v^{(0)}$, or more generally, of \hat{Q}_{M} , which appear in (5.2), (5.3) above, are readily indicated: we have

$$\langle \hat{Q} \rangle = X \langle \langle v^{(0)} \rangle \rangle$$
, where $\langle v^{(0)} \rangle = \langle \langle v^{(0)} \rangle_{R} \rangle_{S}$. (5.6)

This is different from zero, in scattering, if specular reflection geometry (both on a surface and in a volume) is available. The second-order moments are instructive:

$$\langle \hat{Q}_1 \hat{Q}_2 \rangle = \chi \{\langle v_1^{(0)} v_2^{(0)} \rangle \}$$
, (5.7a)

where now

$$\langle v_1^{(0)} v_2^{(0)} \rangle = \langle \langle \Delta v_1^{(0)} \Delta v_2^{(0)} \rangle_{R} \rangle_{S} + \langle \langle v_1^{(0)} \rangle_{R} \langle v_2^{(0)} \rangle_{R} \rangle_{S}.$$
 (5.7b)

Equation (5.7b) contains both incoherent $(\sim \Delta v^{(0)})$ and coherent terms $(\sim < \Delta v^{(0)}>_R)$, of all orders $(k \ge 1)$ of coupled scattering or radiating elements. The associated covariance of \hat{Q} is

$$K_{\hat{0}} = \mathcal{L}\{\langle v_1^{(0)} v_2^{(0)} \rangle - \langle v_1^{(0)} \rangle \langle v_2^{(0)} \rangle \},$$
 (5.7c)

which embodies wholly incoherent radiation.

We now distinguish two principal classes of radiation models from which scattered and ambient acoustic (and electromagnetic) noise fields can be constructed. These are (1): the "classical" cases, where the number of "radiation events" (e.g., reradiations or ambient emissions) are the same for each member of the ensemble, so that $\langle v^{(k)} \rangle_R = v^{(k)}$, and $\langle v^{(k)} \rangle_R = v^{(k)}$, and $\langle v^{(k)} \rangle_R = v^{(k)}$, and $\langle v^{(k)} \rangle_R = v^{(k)}$, can be nonvanishing, reflecting the fact that the number of radiation events is variable over the ensemble. The latter is typical of the so-called poisson radiation models [11], developed recently for both scattering [28]-[32], [40] and ambient noise [3]-[6], [52].

For the poisson scattering cases there is the so-called FOM (Faure [28], Ol'shevskii [29], Middleton [31]) model, based on an independent poisson point-process, where

(FOM):
$$v^{(0)} = v^{(1)}$$
; $v^{(k)} = 0$; $k \ge 2$, (5.8)

namely, all multiple scatter is neglected. For this model the coherent and incoherent contributions are found to be linear functionals of

$$\langle v^{(0)} \rangle_{R} = \langle v^{(1)} \rangle_{R}, \quad \text{or} \quad \langle v^{(0)} \rangle = \langle v^{(1)} \rangle \equiv \sigma^{(1)}; \qquad (5.9a)$$

$$(\underline{\text{FOM}}): \quad K_{\Delta v}(0) = \langle v^{(0)}_{1} v^{(0)}_{2} \rangle - \langle v^{(0)}_{1} \rangle \langle v^{(0)}_{2} \rangle$$

$$= \langle \sigma^{(1)} \delta(\underline{z}_{2} - \underline{z}_{1}) \rangle_{S} + \langle \langle v^{(1)}_{1} \rangle_{R} \langle v^{(1)}_{2} \rangle_{R} \rangle_{S} - \langle v^{(1)}_{1} \rangle \langle v^{(1)}_{2} \rangle, \qquad (5.9b)$$

where \underline{z}_1 , \underline{z}_2 embody spatial, temporal and parametric coördinates and $\sigma^{(1)}(\underline{z})$ is the average density of scattering in the illuminated region Λ . The FOM model is quasi-phenomenological and avoids dealing explicitly with the boundary conditions, which are now contained in a (linear) ad hoc response function $Y_0h(t-t_2)$, with "cross-section" Y_0 , developed and

empirically adjusted for the problem at hand [43]-[45].

On the other hand, in the classical approach, the density fluctuations of radiation events are always zero, as noted above, so that (5.7b), $\langle v_1^{(0)} v_1^{(0)} \rangle = \langle v_1^{(0)} v_2^{(0)} \rangle, \text{ is an identity, and (5.7c) describes the covariance of } v_1^{(0)}, \text{ or } \hat{\mathbb{Q}}. \text{ Again, in the usual applications, multiple scatter is neglected } (v_1^{(k)} = 0, k \ge 2). \text{ One has, here for surface scatter cases, specifically}$

$$\gamma_{o} < v^{(0)} >_{R} = \gamma_{o} v^{(0)} = \gamma_{o} v^{(1)} = RS \frac{\omega}{c} \hat{n}_{g} (\hat{j}_{T} - \hat{j}_{R})$$
 (5.10)

under a Kirchoff approximation, [27], [28], where R, S are respectively the (average) plane-wave reflection coefficient and shadowing function, ω (=2 π f) is the angular frequency of the incident radiation, c_0 = phase speed of sound in water, $\hat{n}_g = (\hat{1}_x \zeta_x + \hat{1}_y \zeta_y - \hat{1}_z)/\sqrt{1+\zeta_x^2+\zeta_y^2}$ is the normal to the wave surface $\zeta(\underline{r},t)$, with $\zeta_x = \partial \zeta/\partial x$, etc., and $\hat{1}_T$, $\hat{1}_R$ are unit vectors from source and receiver to the scattering element on ζ . See [25]-[27], [37]-[39] as well as [47]-[51], for detailed development of the classical theory. See Sec. 6 following and Appendix A.2 for further discussion of these poisson field models.

6. Poisson Statistical Models: Non-Gaussian Noise Fields and Received Processes, X

Because of the effectively discrete nature of the inhomogeneity models and the analogous structure of the ambient field sources, the fundamental statistics of both the fields $\alpha(R,t)$ generated in the medium and the received waves, X(t), are poissonian. This is directly the case for the ambient noise and weakly scattered field components, which constitute much of the interference in practical operation. In many cases, i.e., where there may be a few relatively strong sources or scatterers outstanding from the majority, the resultant field and received process X are highly nongaussian, while the weaker background is, as expected, well described by a limiting gaussian process. We note some of the important results below.

6.1 Poisson Fields and Statistics

A (scalar) poisson field $\iota(\underline{R},t)$, here an underwater acoustic field generated by potentially many independent sources, may be represented by

$$\alpha(\underline{R},t) = \int_{\underline{Z}(=\underline{Z}_{R}\times\underline{Z}_{S}\times\underline{Z}_{\Theta})} \underline{X}(\underline{R},t|\underline{z})dN(Z(\underline{z})), \qquad (6.1)$$

where $\underline{z} = \underline{z}_R \times \underline{z}_S \times \underline{\theta}$, in which \underline{z}_R represent radiation epochs, \underline{z}_S are spatial variables, and $\underline{\theta}$ are structural or waveform parameters; $\underline{Z} = \underline{Z}_R \times \underline{Z}_S \times \underline{Z}_{\theta}$ is the associated composite space of these quantities; $\underline{Z}_S \times \underline{Z}_{\theta}$ is the wave field for a single emitting source. Here \underline{d} N is a counting functional, and in general, \underline{d} N, \underline{z}_R , \underline{z}_S , $\underline{\theta}$ are random (process and) variables. This counting functional is represented by

$$dN(\underline{Z}) = \sum_{p=1}^{p} \delta(\underline{\lambda} - \underline{\lambda}_{p}) \delta(\underline{\theta} - \underline{\theta}_{p}) \delta(t - t_{p}), \qquad (6.2)$$

for P discrete sources potentially emitting, or reradiating, at times \mathbf{t}_n in region $\Lambda,$ where

$$dN(\underline{Z}) = dN(\Lambda = p; t_p; \underline{\theta}) = (\underline{dZ})^p e^{-\underline{dZ}}/p!, \quad p \ge 0$$
 (6.2a)

with $\underline{dZ} = \sigma_s d\Lambda \underline{d\theta}' dt' w_1(\underline{\theta}') w_1(t')$, etc., in which t' is a random emission time or epoch. For poisson fields, dN is a poisson process obeying (6.2a) such that for dn = dN - $\langle dN \rangle$,

$$\langle dn_1...dn_Q \rangle = \rho(z_1) dz_1 \prod_{j=1}^{Q-1} \delta(z_{j+1}-z_j) dz_{j+1}; \langle dn_1 \rangle = 0;$$
 (6.3)

e.g.,

$$\langle dn_1 dn_2 \rangle = \rho(\underline{z}_1) \delta(\underline{z}_2 - \underline{z}_1) \underline{d}\underline{z}_1 \underline{d}\underline{z}_2, \qquad (6.3a)$$

$$\langle x(R,t) \rangle = \int \mathcal{Z}(R,t|z) \rho(z) dz;$$
 (6.4a)

$$K_{z}(\underline{R}_{1},t_{1};\underline{R}_{2},t_{2}) = \int \underline{\mathcal{X}}(\underline{R}_{1},t_{1}|\underline{z})\mathcal{X}(\underline{R}_{2},t_{2}|\underline{z})\varrho(\underline{z})d\underline{z}, \qquad \qquad \begin{cases} (6.4b) \end{cases}$$

with analogous results for the higher moments. Since $X = \hat{R}\alpha$, cf. (2.4b), (5.1), the corresponding results for the received process X are

$$\langle X(t) \rangle = \int_{\mathcal{Z}} L(t|z) \rho(z) dz, \quad L = \hat{R} \mathcal{L}$$

$$(6.5a)$$

$$\langle X(t) \rangle = \int_{\mathcal{Z}} L(t|z) \rho(z) dz, \quad L = \hat{R} \mathcal{L}$$

$$(6.5b)$$

$$K_{\chi}(t_1, t_2) = \int_{\underline{Z}} L(t_1|\underline{z}) L(t_2|\underline{z}) \rho(\underline{z}) d\underline{z}.$$
 (6.5b)

The $Q^{\frac{th}{-}}$ order characteristic function (c.f.) is readily found to be ([10], [31], Part I)

log
$$F_0(i\xi_1,\ldots,i\xi_0|\alpha_1,\ldots,\alpha_n \text{ or } X_1,\ldots,X_0) =$$

$$<\int \rho(\underline{z}',z)[\exp\{i\sum_{\ell=1}^{Q}\xi_{\ell}(\mathbf{Z}_{\ell} \text{ or } L_{\ell})\} -1] dz>_{\underline{z}'},$$
 (6.6)

where $\underline{z}=(\underline{z}',z)$ and $z=t/\overline{1}_S$ is a normalized time, with $\mathbf{L}_1=\mathbf{L}(\underline{R}_1,t_1|\underline{z})$, $L_1=L(t_1|\underline{z})$, etc. [The limiting gaussian cases follow formally from (6.6) by expanding the exponent through $O(\xi_k\xi_\ell)$ and dropping higherorder terms.] The associated $Q^{\frac{th}{t}}$ -order probability density function (pdf) is formally, with $y=\mathbf{L}$ or L:

$$W_{J}(y_{1},...,y_{Q}) = \int_{-\infty}^{\infty} \exp \left\{ i \sum_{\ell} \xi_{\ell} y_{\ell} \right\} F_{n}(i \underline{\xi} | \underline{y}) \frac{d\xi}{(2\pi)^{Q}}; \quad \underline{d\xi} = d\xi_{1}...d\xi_{Q}, \text{ etc.}$$

$$(6.7)$$

Evaluating (6.7) based on (6.6) is generally nontrivial. Only the cases n=1,2 appear analytically manageable, but fortunately $w_1(y)$ is sufficient to give us useful results in signal processing applications, as we shall observe in Part II following. [The case Q=2, e.g., $w_2(y_1,y_2)$ allows us to extend the "classical" theory of signals and noise through (zero-memory) nonlinear devices to include specific nongaussian noise processes and fields, as recent work has shown [10].] We remark that our ability to construct the field covariance (6.4b) allows us to estimate the degree of spatial processing gain $(1 \le M' \le M)$ attainable when the

field is nonuniform in space. Here M' is the estimated number (not necessarily integral) of effectively independent sensors, out of a total M, cf. comments in Sec. 1 above, and Refs. [3]-[5]; also Sec. 8.3 ff.

As a specific example, for the many "far-field" situations where it is meaningful to talk of beams and beam patterns, cf. Sec. 4.2, the basic (complex) field $\boldsymbol{\mathcal{L}}$ can usually be described (in the far-zone) by an expression of the form

$$\hat{Z} = \frac{\hat{S}_{oT}}{\lambda^{\gamma}} \int_{-\infty}^{\infty} \hat{S}_{T}(f|z_{o}) Q_{T}(\lambda|f) e^{i\omega\mu_{d}(t-\varepsilon-|R_{o}-R|/c_{o})} df, \qquad (6.8)$$

where now **%** in (6.1), (6.4), (6.6) is **%** = Re $\hat{\mathbf{x}}$. Here the typical source is treated as essentially a point source, with a non-uniform beam pattern Q_T , Fig. 4.1, and $\lambda = |\underline{R}_0 - \underline{R}|/c_0$, with $\lambda_m = |\underline{R}_0 - \underline{r}_m|/c_0$ the distance of the $m + \frac{th}{c}$ array element from the typical source, and $\gamma > 0$. Here, also,

$$\hat{S}_{T}(f|z) = \left(\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} a(t-\tau)s_{T}^{\dagger}(\tau)e^{i\phi_{a}(t-\tau)}d\tau\right)e^{-i\omega t}dt, \quad \omega = 2\pi f, \quad (6.9)$$

represents the source waveform s_T^* with a fading mechanism, ae t_d^* ; μ_d ($t_d^* = 1 + \epsilon_d$) = doppler coefficient in (6.8), while $z = (t_d^* - 1) + \epsilon_d$ is a (normalized) time, cf. (6.6).

From the above we can show, for instance, that the covariance of the interference output of the receiving array, i.e., $X(t) = \hat{R}\alpha$, (4.12), becomes, for these sources narrow-band about frequency f_0

$$| K_{1}(\tau)_{R} = \frac{1}{2} B^{2} A_{(A \text{ or } B)} < \frac{1}{R^{2\gamma}} > \text{Re}[M_{a}(\tau)D_{1}(\omega_{o}; \hat{K}_{o}(\tau) < \hat{Q}_{R}(\underline{\psi}_{o} - \underline{\psi}_{oR}, f_{o})|^{2} > \hat{\underline{1}}_{o}],$$

$$(6.10)$$

where $^*M_a(\tau)$ = complex covariance of the fading, ae $^{i\phi}a$; $D_1(\omega_0\tau)$ = $<e^{i\omega_0\mu\tau}>_{\epsilon_d}$ is the doppler "smear" factor, and

^{*} $\hat{\kappa}_0$ is the typical interfering source covariance (of s_{τ})

represents the averaged beam pattern, over all the different angles (θ_0,ϕ_0) of arrival of the various interfering source wavefronts; $A_m = (\text{complex})$ element weighting, A = normalization factor. for example, for an equally weighted vertical array (4.14a) when $w_1(\theta_0,\phi_0) = 1/\pi^2$, $-\pi/2 < \theta_0 \le \pi/2$; $0 \le \phi_0 \le \pi$, i.e., (θ_0,ϕ_0) are each uniformly distributed, for the mm' element in (6.10a) we get

$$\langle Q_{R}^{(m)} Q_{R}^{(m')*} \rangle_{\hat{\mathbf{J}}_{0}} = e^{2\pi i \Delta_{mm'} \sin \theta_{0} R} J_{0}(2\pi \Delta_{mm'}); \quad \Delta_{mm'} = (m-m')\Delta \ell / \lambda_{0}, \quad (6.10b)$$

where $\Delta\ell$ is the interelement spacing and λ_0 (= $2\pi c_0/\omega_0$). Generally, the random angles of arrival "smear" the beams. Moreover, Eq. (6.10b) in (6.10a), (6.10) shows at once the (statistical) nonuniformity of these interference fields, as seen over a typical array, when the array is large enough, e.g., $2M\Delta\ell/\lambda_0 = L/\lambda_0 = O(1 \text{ or more})$ here, cf. $J_0(2\pi\Delta_{mm^1}/\lambda_0)$ in (6.10c), etc.: the field covariance as sampled by the array elements is clearly quite nonuniform. [We shall return to this point in Sec. 7.3 ff.]

6.2 Non-Gaussian Process PDF's

As we have emphasized in previous work [4], [6]-[8], [10], [13], [15], [16], [52], involving EMI interference, and analogously here for underwater acoustic applications, most noise environments can be characterized in three main classes of nongaussian process or field:* Class A, or "coherent" noise, producing negligible transients in a typical receiver; Class B, "impulsive," i.e., broadband vis-à-vis the receiver, generating transients; and Class C = Class A + Class B, in various combinations. Typically, Class A--for example, reverberation, ship noise, etc.--or Class B (biologic (shrimp), ambient ice, microseisms, impulsive ship noise, etc.), predominate in most instances.

^{*}An important exception is the case of one (or a small number) of interfering signal sources of known emission behavior, whose deleterious effects can often be removed by selective temporal (= frequency) or spatial deletion.

For our subsequent detection and estimation algorithms (Part II ff.), we shall need a more sophisticated statistic than just the process or field covariance functions, although these are always useful. Specifically, we shall need the (first-order) probability density of the (total) interference, X(t), $w_1(x|H_0)$, as obtained from (6.7), since we are going to sample in such a way as to produce independent <u>noise</u> samples, cf. Sec. 7 ff. For this we set Q=1 in (6.6) and carry out the resulting evaluations, along the lines described in [6], [8]. The result, finally, for Class A interference is found to be, typically, at the $\frac{th}{t}$ array element now

$$F_{1}(i\xi|X^{(m)})_{A+G} = e^{-\sigma_{G}^{(m)^{2}}\xi^{2/2-A_{A}}} e^{A_{A}e^{-\xi^{2}\Omega_{2A}^{(m)}/2}} [1 + \Psi_{1}^{(m)}(\xi^{2})_{A}],$$
(6.11)

When these sources are not so widely distributed, $\psi_1 \doteq 0$, and we have the more familiar "canonical" and approximately canonical Class A cases. For these latter—the pdf associated with (6.11) is

$$w_{1}(x_{m,n}|H_{0})_{A+G} = e^{-A_{A}} \sum_{p=0}^{\infty} \frac{A_{A}^{p} e^{-x_{mn}^{2}/4\sigma_{pA}^{2}}}{p! \sqrt{4\pi\sigma_{pA}^{2}}},$$
 (6.12)

where $\Gamma_A' = \sigma_G^2/\Omega_{2A}$; $2\sigma_{pA}^2 = (p/A_A + \Gamma_A')/(1 + \Gamma_A')$, and $x_{m,n} = x_{m,n}/\sqrt{\Omega_{2A}(1 + \Gamma_A')}$ is the normalized input data sample, at array element m and at time t_n .

(For the "anatomy" of the noise parameters Γ_A' , Ω_{2A} , see [7], [8].) Class B interference is similarly developed: we find that [6]

$$w_{1}(x_{m,n}|H_{0})_{B+G} = \frac{e^{-x_{min}^{2}/\Omega}}{\pi\sqrt{\Omega}} \sum_{p=0}^{\infty} \frac{(-1)^{p}}{p!} A_{\alpha}^{p} \Gamma(\frac{p\alpha+1}{2}) {}_{1}F_{1}(-\frac{p\alpha}{2}; \frac{1}{2}; \frac{x_{m,n}^{2}}{\Omega}).$$
(6.13)

Here $\alpha = \frac{2-\mu}{\gamma}$ is a parameter based on source distribution $(-\lambda^{-\mu})$ and propagation law $(-\lambda^{-\gamma})$, while Ω is a normalization factor [chosen to cut off the tails of (6.13) at some very small residual probability and insure that $\int w_1 dx = 1$: (6.13) does not support a finite intensity, e.g., $\overline{x^2} \to \infty$, unless the pdf is suitably truncated.]

Finally, we emphasize again the canonical nature of these nongaussian models: the <u>forms</u> of the pdf's are independent of the particular noise mechanisms involved. Of course, for specific applications, we must estimate the relevant parameter values and insert them appropriately in the chosen algorithm. Thus, we may say that signal processors based on these noise models are <u>parametrically adaptive</u>, beased on the above statistical-physical models of the acoustic noise or interference environment.

6.3 Summary Remarks (Part I)

In the preceding sections (Part I) we have outlined a formal apparatus, illustrated by some simple examples and specific results. This apparatus provides a general approach for obtaining the needed noise, signal fields, and received process models, which constitute the inputs to our subsequent signal processors, namely, signal detectors and estimators. Key features of these inputs which must be taken into account are (i) the generally nongaussian character of the accompanying noise, or additional acoustic interference from localized sources; (ii) the spatial, as well as temporal character of the noise and signals; (iii) the salient features of propagation in a typically inhomogeneous medium which includes discrete, continuous, and distributed inhomogeneities; and (iv) the spatial coupling to the medium itself. The controlling factors, as always, are 'timately physical. These, in turn, may be expected to impact strongly upon the

design and operation of the subsequent signal processing, as we shall see in Part II following. Finally, but not least, we stress again the direct analogy between the underwater acoustic field and process models outlined here and corresponding EMI models developed for telecommunication applications: the formal structures, and ambient noise models (neglecting multiple scatter) are the same, as are the basic processing concepts, so that results from the latter discipline can be translated to the former, with appropriate attention to the governing physics. In any case, our approach emphasizes the concept of the generalized channel, in the manner of Fig. 2.1 above, where in addition to propagation questions, we include the necessarily coupled signal processing ones, centered on the critical limiting threshold cases.

Part II. Acoustic Threshold Signal Processing

7. Threshold Signal Detection

We are now ready to outline the aforementioned canonical approaches to signal processing, when the desired signal is weak compared to the noise background. Here, of course, the desired signal (and undesired noise) are the received, scattered and/or ambient acoustic fields, now sampled by our distributed arrays (cf. Sec. 4.2). Our principal tasks are to detect a desired signal, and then, having determined its existence in the accompanying noise, to extract some one or more desired attributes of the signal, namely, to perform so-called "signal estimation."

As before [7], [15], [16], we focus our attention primarily on the "on-off" threshold, or weak-signal detection situation, involving now nonuniform interference fields (I), where the decision process is the hypothesis test: H_1 :S \oplus I vs. H_0 :I alone. A canonically optimum theory is again possible: canonical in the sense that the formal results for detector structures and performance are independent of any detailed physical mechanism, as well as of specific signal waveforms and noise statistics [16]. Although the results are no longer optimum for stronger signals, they are generally absolutely better than for weak signals, which in turn is normally satisfactory for applications. Moreover, an optimal

theory suggests reasonable suboptimum procedures which can have the advantage of structural and operational simplicity, as well as performance close to optimum [15], [16].

Three modes of detection are noted: (i) <u>coherent</u>, when signal epoch is known precisely at the receiver; (ii) <u>incoherent</u>, when signal epoch is unknown; and (iii) <u>composite</u>, in which a sum of modes (i) and (ii) is employed, to take advantage of any signal epoch information available [16]. Since weak signals are postulated, large processing gains (here <u>space</u>-time-bandwidth products J) are needed to achieve acceptably small probabilities of decision error. This ensures that the detection algorithm in question will be asymptotically optimum (A0) (as $J \rightarrow \infty$, or J >> 1 practically), and normally distributed under H_0 , H_1 provided a suitable <u>bias</u> term B_J^* (independent of the data and determined a priori) is employed. (The importance of the proper bias, generally, is stressed in [16].) These results enable us to calculate performance directly in canonical form.

For specific applications we must, of course, "calibrate" our canonical algorithms to the particular interference environment. This requires establishing: (1) the Class (A, B, or C) of interference, and (2) estimating the relevant parameters of the associated probability densities $\mathbf{w}_1(\mathbf{x})_A$, $\mathbf{w}_1(\mathbf{x})_B$, etc., [15], [54]. This can be done if we can construct the EMI scenario, for example [4], [8], or directly by empirical observation [6], [7]. Thus, these (detection) systems are adaptive and include adaptive beam-forming here. Their often considerable improvement over conventional (i.e., "matched-filter" or correlation) detectors stems from this essential feature, [7], [16].

In this necessarily concise overview we confine our effort to an outline of some of the principal results of the extensions of our earlier work to include spatial processing [3], [4]. The present section is organized as follows: Sections 7.1-7.4 provide examples of general and specific space-time processing algorithms, the latter for independent sampling; [Section 6 earlier presents some results for nongaussian field models]. Section 8 ff. treats performance, in both optimum and suboptimum cases. Section 8.3 completes our detection treatment here, with a brief discussion of the principal assumptions and conditions.

7.1 Canonical Space-Time Threshold Algorithms

To obtain the desired optimum (binary) decision algorithms ℓ_J^* we form first, as usual [20], [54], the generalized likelihood ratio, now extended to include spatial as well as temporal sampling, e.g.,

$$\ell_{\uparrow}^{\star} \equiv \log \Lambda_{\uparrow}(x|\theta) = g_{\uparrow}^{\star}(x|\theta) + t_{\uparrow}^{\star}(x|\theta), \qquad (7.1)$$

where J = MN, with N = number of temporal samples of the received data $\underline{x} = (x_1 \dots x_{j=mn} \dots x_J)$, at each of the M sensors, e.g., $\underline{x}_m = (x_{m1}, \dots, x_{mn=j}, \dots x_{mN})$, with $j = 1, \dots, mn, \dots, J = MN$, alternatively, over all sensors. Here θ denotes an input signal-to-noise ratio; g_J^* , t_J^* are respectively the desired algorithm and a remainder series, which vanishes prob. 1 under H_0 , H_1 if certain conditions are satisfied [cf. Sec. 7.4 ff.].

For these threshold cases $\log \Lambda_J$ is expanded about $\theta=0$, with one or two terms in the data (\underline{x}) retained, depending on the mode of reception. The required bias term, $B_J^*(\theta)$, is obtained from the average over $\{\underline{x}\}$ with respect to the null hypothesis H_0 of the next nonvanishing term in the g_J^* -expansion. Thus, we write formally

$$\mathcal{L}_{J}^{*} = g_{J}^{*} = \theta F_{1}(\underline{x})_{J} + \frac{\theta^{2}}{2!} F_{2}(\underline{x})_{J} + O(\theta^{3}, \theta^{4}), \qquad (7.2)$$

where for the coherent and incoherent detection modes we have

$$g_{J}^{\star}(\underline{x})_{coh} = \theta F_{1}(\underline{x})_{J} + B_{J}^{\star}(\theta^{2})_{coh};$$
 $B_{J-coh}^{\star} = \frac{\theta^{2}}{2!} \langle F_{2}(\underline{x})_{J} \rangle_{H_{0}} + \log \mu;$ (7.3a)

$$g_{J,inc}^{*} = \frac{\theta^{2}}{2!} F_{2}(\underline{x})_{J} + B_{J}^{*}(\theta^{4})_{inc}; \quad B_{J-inc}^{*} = \langle 0(\theta^{4};\underline{x}) \rangle_{H_{0}} + \log \mu.$$
 (7.3b)

The specific structure of F_1 , F_2 depends, data-wise, on the basic pdf $w_J(\underline{x}|\theta=0)$ of the interference alone, and on the desired signal waveform structure. Here $\mu=p/q$, p+q=1, (p,q>0), where p,q are respectively the a priori probabilities that the data sample \underline{x} contains (or does not contain) a signal. The decision process (§) is

decide H₁ if
$$g_{J}^{*} \ge \log K$$
 or decide H₀ if $g_{J}^{*} < \log K$ (7.3c)

where K is a threshold in the usual way.

Both because of much needed technical simplicity, and the fact that it can be shown [4] that comparatively little further improvement in performance is obtained by fully correlated sampling vis-à-vis (non-sparse) independent (space and time) sampling, we focus henceforth on this latter case. Moreover, although the theory has been developed for the general case of nonstationary and inhomogeneous fields [4], we shall describe results here only for the (often) applicable local stationary, homogeneous régimes.

7.2 LOBD Algorithms

Locally optimum Bayes detector (LOBD) algorithms for coherent and incoherent detection are found under the above detection modes (7.3a), (7.3b) to be explicitly [4]:

A. Coherent Detection

$$g_{J}^{\star}(\underline{x})_{coh} = B_{J-coh}^{\star} - \sum_{m=1}^{M} \sum_{n=1}^{N} \ell(x_{m,n}) < \theta_{m,n} > , \qquad J = MN$$
 (7.4)

with

$$\langle \theta_{m,n} \rangle = \langle a_{on} s_n^{(m)} \rangle; \quad \ell(x_{m,n}) = \frac{d}{dx} \log w_1(x|H_0) \Big|_{x=x_{m,n}} \quad (= \ell_{m,n}).$$
 (7.4a)

B. Incoherent Detection

$$g_{\tilde{J}}^{\star}(\underline{x})_{inc} = B_{\tilde{J}-inc}^{\star} + \frac{1}{2!} \sum_{m,m'}^{M} \sum_{n,n'}^{N} \{\ell_{m,n}\ell_{m',n'} + \ell'_{m,n}\delta_{mm'}\delta_{nn'}\} < \theta_{m,n}\theta_{m',n'} >$$

$$(7.5)$$

where now

$$<\theta_{m,n}\theta_{m',n'}> = \overline{a_0^2} \hat{m}_{|n-n'|}\rho_{|n-n'|}^{(mm')}$$
 (7.5a)

with

$$\widehat{m}_{|n-n'|} \equiv \overline{a_{on}a_{on'}} / \overline{a_o^2} ; \qquad \rho_{|n-n'|}^{(m-m')} \equiv \langle s_n^{(m)} s_n^{(m')} \rangle , \qquad (7.5b)$$

and

$$\ell_{m,n} = \frac{d\ell}{dx} |_{x=x_{m,n}}$$
, cf. (7.4a).

In the above, $s_n^{(m)} = s^{(m)}(t_n)$ is the normalized signal waveform at time t_n at the $\frac{th}{n}$ -sensor, with $\langle s_n^{(m)} \rangle = 1$; $a_{on}^{(m)} = a_{on} = a_{o}(t_n)$ here (signal field uniform over the array), with $a_{on} = A_o(t_n)/\sqrt{2\psi}$, in which $\psi = I = (I' + N)$ is the sum of the intensities (at the receiving aperture) of the nongaussian interference I' and gaussian noise N. (Here $\psi_n^{(m)} \to \psi$, because of the postulated homogeneity and stationarity, cf. above.) The normalized data samples are $\chi = \chi_{j=m,n}/\sqrt{\psi}$, with the desired signal vector $S = \{S_o^{(m)}\} = \{a_{on}^{(m)}s_n^{(m)}, \sqrt{\psi}\}$. Here w_1 is the first-order pdf of the total interference: because of independent sampling $w_j = \frac{\pi}{m}$, $w_1(x_m,n)$.

7.3 Signal Structures for Coherent and Incoherent Reception

Specifically, we note also that in realistic situations involving fading and doppler uncertainties the desired signal structure is

$$s_{n}^{(m)} = \sqrt{2} \cos\{(\omega_{0} + \omega_{d}) (t_{n} - \varepsilon_{0}) \omega_{0} \Delta \tau_{m} - \phi_{n}\}$$

$$t_{n} = n \Delta t$$
(7.6)

where ω_d = random doppler shift, $\varepsilon = \varepsilon_0$ = signal epoch which is a priori known in <u>coherent detection</u>; ϕ = possible phase modulation, for the typical narrow-band signals used in most acoustic environments. The net path delay, $\Delta \tau_m$, from the m + 1 - element to some selected reference point associated with the array, is given by

$$\Delta \tau_{m} = (\hat{\mathbf{1}}_{0} - \hat{\mathbf{1}}_{0R}) \cdot \mathbf{r}_{m}/c_{0} = (\underbrace{\vee_{0} - \vee_{0R}}) \cdot \mathbf{r}_{m}/f_{0}, \qquad (\underbrace{\vee} = \hat{\mathbf{1}}/\lambda_{0}; c_{0} = \lambda_{0}f_{0})$$
(7.7)

where $\hat{\underline{1}}_0$, $\hat{\underline{1}}_{0R}$ are respectively the unit vectors in the direction of the (incoming) signal wavefront, and the main axis of the beam formed by

the equivalent spatial alignment of the M elements; \underline{r}_m is the (vector) distance from the \underline{m}^{th} -element in the array to the reference point. Thus, when the beam is steered to the signal direction, $\hat{i}_{OR} = \hat{i}_{O}$, cf. Fig. 4.1, and $\Delta \tau_m = 0$, so that $s_n^{(m)} = s_n$: the sampled signal component is now independent of m in (7.4), (7.4a) above, for coherent detection.

For the <u>incoherent cases</u> (7.5), (7.5a,b), (where ε_0 is now random), we have, for a fixed signal wavefront direction,

$$\begin{bmatrix}
\rho(mm') \\
|n-n'| = e
\end{bmatrix} = e^{-\left[\Delta\omega_{d}(n-n')\Delta t\right]^{2}/2} \cos\left[\omega_{o}(n-n')\Delta t + k_{o}(\hat{j}_{o}-\hat{j}_{oR})\cdot\Delta r_{mm'}\right],$$

$$k_{o} = 2\pi/\lambda_{o} = \omega_{o}/c_{o}$$
(7.8)

with $\Delta r_{mm'} = r_m - r_m'$, and there is no degradation of the beam pattern. This signal correlation function, $\rho_{|n-n'|}^{(mm')}$, is clearly maximized (when summed over m,m', in (7.5), for example) when the receiving beam is pointed at the normal $(\hat{1}_0)$ to the incoming signal wavefront, e.g., $\hat{1}_0 = \hat{1}_{0R}$. Thus, $\rho_{|n-n'|}^{(mm')} = \rho_{|n-n'|}$, also. Alternatively, when the desired signal wavefront is perturbed, arriving from various different directions, beam structure is destroyed and $\rho_{|n-n'|}^{(mm')}$ is much reduced, with a consequent serious degradation of performance [cf. (6.10b) and Sec. 3 ff.]. Doppler smearing $(\Delta \omega_d \geqslant E = 0)$, of course, always reduces the wavefront (and waveform) coherence of the desired signal [cf. (7.6), (7.8)], similarly degrading performance, as one would expect [cf. Sec. 6.1 above].

Finally, we note that because of the postulate of independent sampling, space and time processing are interchangeable operations, cf. (7.4), (7.5): we can process over the array (m=1,...M) at any given instant (t_n) , or equivalently, at each array element (m), over time (n=1,...N). The choice is a matter of technical convenience. In addition, we observe that for coherent detection, (7.4), (7.6), (7.7), single beams are formed $(\sim s_n^{(m)})$, while in incoherent detection product beams are generated (m,m') in (7.5).

7.4 The Bias and Asymptotic Optimality (AO)

As explained above, cf. (7.3a,b), the bias terms for g_{j}^{*} are found here to be

$$\hat{B}_{J-coh}^{*} = -\frac{L^{(2)}}{2} \sum_{m=1}^{M} \sum_{n=1}^{N} \langle a_{on}^{(m)} \rangle^{2} \langle s_{n}^{(m)} \rangle^{2}$$
(7.9)

generally. With fading and doppler smear (7.9) becomes

$$B_{J-coh}^{*} = \log \mu \equiv \hat{B}_{J-coh}^{*} \stackrel{:}{=} -a_{O}^{2} (1-\eta)MNL^{(2)}H_{1}(N\Delta t\Delta \omega_{d}), \qquad (7.9a)$$

where $0 \le 1-\eta \le a_0 >^2/(a_0^2) < 1$ is a fading parameter (no fading: $\eta = 0$; deep fading: $\eta = 1$), and

$$H_1(x) = \sqrt{\pi}\Theta(x)/2x$$
, with $\Theta(x) = (2/\sqrt{\pi})\int_0^x e^{-t^2} dt = \text{erf } x$, (7.9b)

represents the effects of doppler smear.

Similarly, for the incoherent cases one gets

$$B_{J-inc}^{*} = -\frac{1}{8} \sum_{m,m'}^{M} \sum_{n,n'}^{N} \{(L^{(4)}_{-2L}^{(2)})^{2}\}_{\delta_{mm'},\delta_{nn'}+2L^{(2)}}^{(2)}\}_{\rho_{[n-n']}^{(mm')}}^{(m,m')}^{2}$$

$$\cdot \langle a_{0}^{(m)}a_{0}^{(m')}\rangle \langle s_{n}^{(m)}s_{n}^{(m')}\rangle$$
(7.10)

which reduces for rapid (one-sided) fading and negligible doppler to

$$B_{J-inc}^{*} - \log \mu = \hat{B}_{J-inc}^{*} = -a_{o}^{-2} \left\{ \frac{MN(1-\eta)L(2)}{2} \right\}^{2} \qquad (7.10a)$$

(For no or slow fading one sets $\eta=0$ in (7.10a).) The result (7.10a) assumes N >> 1 (M \geqslant 1) and the usual nongaussian noise situation L⁽⁴⁾ \simeq 2L^{(2)²}, where specifically [4], [16]

$$L^{(2)} = \langle \ell^2 \rangle_{H_0}, \text{ all m,n,} = \int_{-\infty}^{\infty} (w_1'/w_1)^2 w_1 dx, > 0, \text{ cf. } (7.4a);$$

$$L^{(4)} = \int_{-\infty}^{\infty} (w_1''/w_1)^2 w_1 dx \ (>0).$$
(7.11b)

Thus, $L^{(2)}$, $L^{(4)}$ are statistics of the (total) interference, I, including the gaussian noise component.

For <u>asymptotic optimality</u>, ("A0" as $J \rightarrow \infty$, or J >> 1) a sufficient condition--always fulfilled for small (> 0) input signals, e.g., $1 >> \overline{a_0^2} > 0$ -- is that

$$\operatorname{var}_{H_0} g_{J}^{*} \equiv \langle (g_{J}^{*})^2 \rangle_{0} - \langle g_{J}^{*} \rangle_{0}^{2} \equiv (\sigma_{0J}^{*})^2 = -2\hat{B}_{J}^{*} \quad (= \operatorname{var}_{H_1} g_{J}^{*}). \quad (7.12)$$

As noted above, this permits us to terminate the expansion (7.1) in the form (7.2), for large J(=MN), and still ensure the desired small probabilities of decision error in the decision process (7.3c). [An equivalent (sufficient) condition for AO is that g_{0}^{*} , as $J \rightarrow \infty$, is gaussian, with variance $(\sigma_{0J}^{*})^{2}$ and means $\log \mu \mp (\sigma_{0J}^{*})^{2}/2$ under H_{0} , H_{1} respectively, cf. Sec. 2B of [16].] (In most earlier work the proper bias was omitted, cf. Secs. IIb, IV, VA of [16], so that the resulting detection algorithms are neither locally optimum nor AO.)

7.5 Suboptimum Algorithms

Suboptimum algorithms are handled in the same general fashion. Instead of the quite complex characteristics (${}^{\sim}\ell_{m,n}$) in g_J^{\star} , (7.4), (7.5) one can use much simpler forms: for the simple correlator one sets $\ell_{m,n} \to x_{m,n}$, and for the hard-limiter correlator, $\ell_{m,n} \to sgn(x_{m,n})$ (with the $\ell_{m,n}$ term omitted in (7.5)). The former is well-known to be (threshold) optimum generally in gauss noise, $w_1(x) = e^{-x^2/2} / \sqrt{2\pi}$, while the latter is threshold optimum in Laplace noise, $w_1(x) = (\sqrt{2})^{-1} \exp(-\sqrt{2}|x|)$, $< x^2 > = 1$; < x > = 0, cf. Sec. III of [11].

8. Threshold Detector Performance

Because of the large samples involved (J=MN>>1) and their independence, the test statistic g_3^* is asymptotically normally distributed, under both H_0 and H_1 with var $_{H_0}$ g_3^* , provided the input signal is suitably small [cf. (7.12) and II,B of [16], etc.]. This permits a direct calculation of performance. For the important detection cases here, we have the familiar Neyman-Pearson (N.P.) form [15] of the probability of correctly detecting the presence of a desired signal:

$$P_{D}^{(*)} \simeq \frac{p}{2} \{1 + \Theta \left[\frac{\sigma_{OJ}^{(*)}}{\sqrt{2}} - \Theta^{-1} (1 - 2\alpha_{F}^{(*)}) \right] \},$$
 (8.1)

where $\alpha_F^{(*)}$ is the (conditional) false alarm probability of incorrectly deciding the signal is present when only noise occurs, again $\Theta(x) = \text{erf } x$, cf. (7.9a), and

$$\sigma_{\text{OJ}}^{(\star)} = \sqrt{\text{var}_{H_0} g_{\text{J}}^{(\star)}}. \qquad (8.2)$$

For suboptimum detectors we drop the superscript *. Explicitly, $\alpha_{\mathsf{F}}^{(\star)}$ is given by

$$\alpha_{\mathsf{F}}^{(\star)} \simeq \frac{1}{2} \left\{ 1 - \Theta \left[\frac{\sigma_{\mathsf{oJ}}^{(\star)}}{2\sqrt{2}} + \frac{\log(\mathsf{K}/\mu)}{\sqrt{2}} \right] \right\}; \tag{8.3}$$

for μ , cf. Eq. (7.3) et seq. In the usual applications, μ = 1, i.e., $p = q = \frac{1}{2}$. (In communication applications, when the decision criterion in the so-called Ideal Observer [15], [16], μ = 1 also, i.e., $p = q = \frac{1}{2}$ and one has a symmetric "on-off" channel for which the <u>error probability</u> (per decision or per baud, symbol, etc.) is now

$$P_e^{(\star)} = pp_e^{(\star)} \approx \frac{1}{2} \left\{ 1 - \Theta \frac{\sigma_{oJ}^{(\star)}}{2\sqrt{2}} \right\} [\mu = 1; K = 1].$$
 (8.4)

8.1 Performance Examples

The key element in measuring performance is clearly $\sigma_{oJ}^{(*)}$, cf. (8.1)-(8.4). We list below a variety of specific forms for $\sigma_{oJ}^{(*)}$:

A. Coherent Detection:
$$(\sigma_{oJ}^{*})_{coh}^{2} = 2a_{o}^{2}(1-n)MNL^{(2)}H_{1}(N\Delta t\Delta \omega_{d}) = 2a_{o}^{2}\pi_{coh}^{*},$$
(8.5)

cf. [4], [15], [16].

B. Incoherent Detection [4]:
$$(\sigma_{OJ}^*)_{inc}^2 = 2a_O^2 \left[\frac{MN}{8} L^{(4)}_{+2L}^{(2)^2} \left(\frac{110}{N} - 11\right)\right] = 2a_O^2 \prod_{inc}^*$$
, (8.6) where Q_N is the signal structure factor, which in these stationary, homogeneous cases becomes explicitly

$$Q_{N} = 1 + \frac{1}{N} \sum_{n,n}^{N} \hat{m}_{|n-n'|}^{2} \hat{p}_{|n-n'|}^{2}; \quad \hat{\vec{u}}_{0} = \hat{\vec{u}}_{0R}, \text{ cf. (7.8) et seq.,}$$
(8.6a)

which is maximized when the beam formed is directed at the desired signal source.

Here $\Pi_{(a_0)}^*$ is the <u>space-time-bandwidth product</u>, and the $\overline{a_0^2}$ factor is defined as the <u>minimum detectable signal</u>, $<(a_0^2)>_{\min}^*$.

Evaluation of Q_{N} for various conditions of doppler "smear" and fading gives results like [4]:

$$(\sigma_{\text{oJ}}^{\star})_{\text{inc}}^{2}$$
 = $\frac{1}{a_{\text{o}}^{2}}$ (MNL(2)/2)² = $(\sigma_{\text{oJ}}^{\star})_{\text{inc}}^{2}$ | 1-sided, no, slow; coherent signal ($\Delta\omega_{\text{d}}^{\pm}0$) (8.7a)

$$(\sigma_{\text{oJ}}^{\star})_{\text{inc}}^{2}$$
 | 1-sided, very rapid; $= \overline{a_{0}^{2}}_{\text{NN}(1-\eta)}^{2}L^{(2)}/2$ (8.7b) coh. sig.

In all cases here we note that $\sigma_{oJ}^* \sim \sqrt{J}$: increasing the "space-time-bandwidth product" by increasing the number of both the independent spatial and temporal samples (M,N) clearly improves performance, from (8.1), (8.4).

It is usually convenient to relate the <u>probability controls</u>, here $p_0^{(*)}$, to the processing gain, $\Pi^{(*)}$, and minimum detectable signal, $a_0^2 > min$, by means of the following relations, obtained from (8.5), (8.6) in (8.7), (8.4), viz.:

$$\langle a_0^2 \rangle_{\text{min-coh}}^{(\star)} = (C_{\text{N.P.}}^{(\star)})^2 / \Pi_{\text{coh}}^{(\star)}; \quad \langle a_0^2 \rangle_{\text{min-inc}}^{(\star)} = C_{\text{N.P.}}^{(\star)} / \Pi_{\text{inc}}^{(\star)}, \quad (8.8)$$

where $\binom{*}{N.P.} = \bigcirc^{-1}(2p_D^{(*)}-1) + \bigcirc^{-1}(1-2p^{(*)}) = \bigcirc^{(*)}_{0J}/\sqrt{2}$ specifically. From (8.8) vs. (8.5), (8.6) we find Π^* at once.

(8.8) vs. (8.5), (8.6) we find Π^* at once. For nongaussian noise fields $L^{(2)}$, $L^{(4)}$ can be quite large $O(10^4,10^8)$, typically, for, say, Class A noise with $A_A = 0.3$, $\Gamma_A = 10^{-5}$. Similar orders of magnitude are noted for typical Class B cases; ([9]; see Figs. 4-7 of [16]). However, for gaussian noise $L^{(2)} = 1$, $L^{(4)} = 2$.

Figure 8.1 shows the probability of detection, $p_D^{(*)}$, versus the minimum detectable signal, $\langle a_0^2 \rangle_{min}^{(*)}$, with processing gain $\pi^{(*)}$ as parameter,

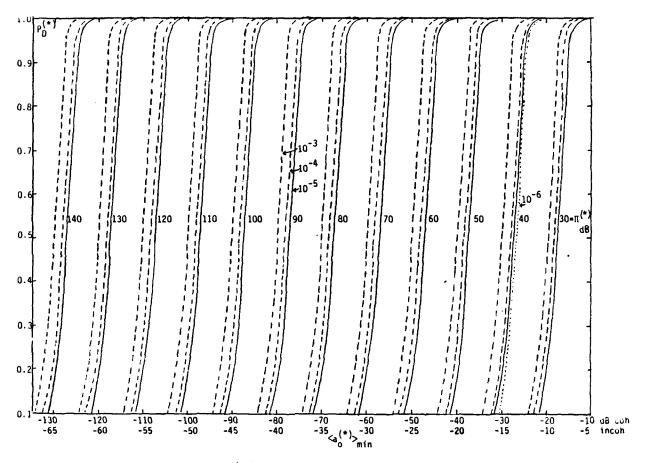


Figure 8.1 Probabilities $P_0^{(*)}$ of detection (8.1) for both coherent and incoherent reception, with processing gain as parameter, for false alarm probability $\alpha_F^{(*)} = 10^{-3}$, 10^{-4} , 10^{-5} .

for various levels of false alarm probability: $\alpha_F^{(*)} = 10^{-3}$, 10^{-4} , 10^{-5} . Clearly, coherent detection is better, by a factor 2 in dB, than incoherent detection with the same processing gain, cf. (8.5) vs. (8.6). For example, with a processing gain of 80 dB, $\alpha_F^* = 10^{-3}$, and a required $p_D^* = 0.80$, we see that $\langle a_0^2 \rangle_{\text{min-coh}}^* = -33$ dB, while $\langle a_0^2 \rangle_{\text{min-inc}}^* = -66$ dB.

8.2 Remarks on Suboptimum Performance

The performance of suboptimum systems is readily obtained canonically (in the threshold régime) if we define a degradation factor $\varphi_{\mbox{d}}$ by the ratio

$$\phi_{\mathbf{d}} = \Pi/\Pi^* \quad , \tag{8.9}$$

Here π is the processing gain for a <u>suboptimum detector</u>, employed against the same (not necessarily gaussian) noise. This factor, ϕ_d , measures by how much the minimum detectable signal for the suboptimum system must be raised vis-a-vis that of the optimum to achieve the same performance, $P_D = P_D^*$ (or $P_D = P_D^*$), viz. $\sigma_{OJ} = \sigma_{OJ}^*$ in (8.1), so that

$$\langle a_0^2 \rangle_{\text{min-coh}}^{\text{min-coh}} = \langle a_0^2 \rangle_{\text{min-coh}} \phi_{\text{d-coh}}; \langle a_0^2 \rangle_{\text{min-inc}}^{\text{min-inc}} = \langle a_0^2 \rangle_{\text{min-inc}} \phi_{\text{d-inc}}.$$
(8.10)

For example, for a coherent detector optimized against gauss noise, $L^{(2)}=1$ in (8.5), so that $\phi_d=(L^{(2)})^{-1}<<1$ measures the degradation when such a detector is used in strongly nongauss noise instead of one properly "matched" to this noise. Thus, if $\Gamma_A'=10^{-5}$, $A_A=0.3$, we have $L_A^{(2)}=50$ dB, Fig. 4, [16], so that the optimum detector can detect a minimum detectable signal 50 dB less than the correlation detector in this noise. In a similar way, two suboptimum detectors can be compared: $\phi^{(1 \text{ vs. } 2)}\equiv \pi^{(1)}/\pi^{(2)}=\phi_d^{(1)}/\phi_d^{(2)}$. The quantity $\sqrt{\phi_d}$ is also seen to be equal to the Asymptotic Relative Efficiency (ARE, $\theta>0$). [See Section IV,B of [16].]

8.3 Discussion: Threshold Detection

In the preceding sections we have sketched some of the main results for the development of optimum threshold detection in nongaussian and nonuniform EMI fields. This extends earlier work, [55]-[57], for example, to include in addition to the nongaussian field, the often important situation where the fields received by an antenna (array) are not uniform over the array elements, basically because of the size of the antenna vis-à-vis the wavelengths of the received fields. In addition, we have outlined briefly various elements in the construction of statistical physical models of the acoustic field, including the needed pdf's.

Our analysis has postulated independent (noise) samples in space and time. The latter are practically achieveable without much difficulty by slightly increasing the sampling period (O(50%)). Independent spatial sampling, however, is not generally possible with non-sparse arrays, because of the non-uniformity of the noise-field and unequal

spacing of the zeros of the correlation functions of this field, cf. (6.10), (6.10a,b). In practice, this means that effectively one may have $1 \le M' < (\cdot) M$ equivalent "independent" spatial samples: some spatial processing improvement over a single element (or single beam, M = 1) is obtainable, but not the full M-factor theoretically achieveable if independent spatial sampling is truly obtained. Nevertheless, for such noise fields as (6.10) (in addition to improving beam resolution), it can be worthwhile to use relatively large arrays $(vs. \lambda_0)$, as M' > 1, and $c. \sigma_{OJ}^*$, $c. \sqrt{M'N'}$, cf. (8.5), (8.6), is larger than $\sigma_{ON}^* = \sqrt{N}$, M = 1.

Of course, algorithms like (7.4), (7.5) are not optimum when the sampling is not independent, but they are comparatively close: preliminary estimates for typical Class A interference indicate that only O(2-3 dB) improvement is theoretically obtainable if all correlations were taken into account, i.e., if one uses the J-joint pdf $w_J(\underline{x})_N$ rather than $j I w_I(x_j)$ for the basic pdf used in the general expansions (7.2)-(7.3b) above. This in one compelling argument for developing "independent-sample" algorithms; another is that there are no tractable, or known $J \frac{th}{t}$ -order models (J > 1.2) for the underlying pdf, w_I .

It is instructive to observe that beam-forming is automatically achieved on setting the relative path delays $\Delta\tau_m$, (7.7), equal to zero, regardless of the particular geometrical locations of the array elements. Thus, we can have arbitrary distributions of elements, but still reconstruct the desired signal wavefront by proper delays (i.e., maximize $\rho^{(mm')}$, cf. (7.8)). Beams formed here in this manner are adaptive, weighted beams. The important trade-offs are coherent space-time signal wavefronts vs. incoherent (uncorrelated) noise field samples, thus increasing the effective statistical sample under H₁ vs. H₀, i.e., lowering the (here) joint probabilities of Type I and II errors, e.g., $\frac{\min}{\delta}$ ($q\alpha*+p\beta*$), $\delta=$ decision rule, viz. (7.3c).

In a broad sense these optimum algorithms seek first, thus, to "match" the receiver to the noise, via $\ell_{m,n}$, cf. (7.4), (7.5), and then to "match" the desired signal to these processed input data, $\ell(x_{m,n})$. It is the "tails" of the noise pdf which are signficiant vis-à-vis gaussian noise: the nongaussian interference falls off much more slowly, so that the transfer characteristic, ℓ , between x-input and ℓ -output includes a

limiting or clipping feature, as well. For this reason, limiters can be good approximations to the actual characteristic, giving performances not many dB worse than the theoretical optimum. Of course, actual differences will depend on the particular EMI field to which the receiver is adapted, i.e., for which the field parameters (Ω_2 , A, Γ' , α , etc.) have been estimated. Finally, it is emphasized that threshold detectors of the above types are useful and give noticeable improvement over conventional systems when the space-time-bandwidth product, $\Pi_J^{(\star)}$, is large, so that the required small probabilities of error $P_e^{(\star)}$ can be obtained. (For a discussion of asymptotic optimality, AO, bounds on < a > o min, etc., as well as some numerical examples, see again [15], [16].)

9. Threshold Signal Estimation

As is well-known, in the Bayes theory [20], [58] estimation is the "twin" of detection: the former aims to determine the particular characteristics of a desired signal, e.g., amplitude, waveform, phase, frequency, etc., once the signal has been detected, while the latter is concerned with the basic question of the desired signal's presence or absence. In a very broad sense "estimation" can be regarded as an extension of detection: both subsume appropriate cost functions for the derivation of optimal algorithms and for the measures and comparisons of performance. In fact, in a more general Bayesian sense detection and estimation are the two, coupled component elements of the composite process of signal extraction: detection and estimation joined together by a suitable cost function, reflecting the frequent situation where it is not completely certain, for the estimation process, that the desired signal is present, which, in turn, results in biased estimators [59].

Here, however, we shall assume that the desired signal is known a priori to be present, and that certain waveform and parameter features of the signal are to be estimated, when the accompanying acoustic interference is highly nongaussian (the "classical" situation of gaussian noise is a special case of our general model). Now the estimators are unbiased. In any case, we may expect the estimation process, particularly optimum estimation (in the Bayes theory [58], [20]), to be closely related

to the detection process, since both are derived from approximate likelihood functions [60], [61]. This will be seen explicitly in the canonical threshold theory outlined below (cf. Sec. 9.1).

From the practical underwater acoustic viewpoint, signal estimation is a common communication requirement: signal amplitude (or level), waveform, frequency, epoch, modulation are each important elements of the reception process, whether it be telecommunications, radar, television (and their analogues in acoustical and optical régimes). Designing receivers for effective threshold performance usually ensures better (but not necessarily optimum) performance at strong-signal levels. As in the detection cases [16], an optimal theory provides limiting estimation algorithms and performance, which are models to be approximated in practice. Also, as in detection, a canonical theory is possible in the weaksignal cases (as long as reasonably large time-bandwith products are permitted, of course: the desired signal must be "extractable" under the observational constraints). Such a canonical theory also provides standards against which practical, suboptimum (i.e., approximate) algorithms can be evaluated and compared [15], [16].

As in threshold detection theory [16] a fundamental problem now in developing effective threshold <u>estimation</u> algorithms is to obtain expressions of limited complexity, which retain their optimum nature when sample size (J) becomes increasingly large. The latter is the case, of course, when the signal is weak, since large effective data samples are required for small expected errors in the resulting estimates—analogous to small probabilities of decision error in optimum (threshold) detection. However, just as in detection [16], without a suitable "bias" term in the extraction algorithm this algorithm demands progressively more terms in its approximative form. This rapidly defeats the key requirement of processing feasibility, particularly for signals and interference of practical use, and destroys analytic tractability as well.

Accordingly, we parallel our preceding summary review of (binary) spacetime threshold detection theory (Secs. 7, 8), and present now a concise summary of optimum and suboptimum threshold signal estimation, for the three generic modes of reception noted earlier for detection: (1) coherent estimation, (2) incoherent estimation, and (3) composite estimation, where a linear combination of the coherent and incoherent algorithms is employed. [Because of spatial limitations we shall confine our explicit illustrations to the simplest situation of amplitude (or scale) estimation, where signal waveform is otherwise known at the receiver. However, the effects of fading, doppler, and other propagation conditions can be included, as in the detection examples above, Secs. 7, 8. The general formalism outlined here, nevertheless, is quite capable of providing explicit algorithms in the more complex situations involving waveform, frequency, phase estimation, etc.] As before, the interference is additive, generally nongaussian, and is either Class A or Class B interference (i.e., respectively coherently or incoherently received in the (linear) front-end stages of a typical narrow band receiver, accompanied by (additive) external and internal gaussian noise [7], [15], [16], [18]. These interference models account for most of the practical cases.

9.1 Canonical Optimum Threshold Estimator Structures; AO/LOBE Forms

A canonical optimum estimator is represented by $\gamma^* = \gamma_{\sigma}^*(\underline{\theta} | \underline{X})$, where $\underline{\theta} = (\theta_1, \dots, \theta_M)$ is the parameter set to be estimated on the basis of \underline{X} ; the received data, σ refers to the pdf of $\underline{\theta}$ (see Appendix A.3). As noted at the beginning of Sec. 9, the critical problem in developing explicit locally optimum threshold estimators (LOBE's) from the general (all signal level) forms (A.3-6), (A.3-8) is to terminate the desired series approximation properly, i.e., without an excessive number of terms and in such a way that these estimators are asymptotically optimum (AO), as well. This situation precisely parallels that for optimum threshold detection, already treated by the author in [15], [20]. In fact, the approach here is to convert the various <u>likelihood functions</u> σF_n , σW_n above, from which the LOBE's are determined, cf. (A.3-4), (A.3-8), into likelihood ratios, equivalent now to a suitable statistical test among two distinguishable hypothesis (H_1, H_2) , etc. Then LeCam's results [60] as applied by Levin [61] are used to obtain at least a sufficient condition of AO for the "LOBD" form from which the resultinf AO LOBE is then constructed, according to the error function (QCF, SCF, etc.) chosen.

A. Equivalent LOBE's for LOBD's

We consider first the SCF case, cf. (A.3-7), with the optimizing rule (A.3-8) for generating the UMLE's, Υ^* . Since the (linear) operation of

differentiation in (A.3-8) is unaffected if we subtract log $F_{J}(X|0)$, we see that (A.3-8) becomes

$$\frac{\partial}{\partial \theta_{\widehat{\mathbf{m}}}} \left\{ \log \sigma(\theta_{\widehat{\mathbf{m}}}) + \ell_{\mathbf{J}}^{(10)}(\underline{\mathbf{X}}|\theta_{\widehat{\mathbf{m}}})_{\mathsf{SCF}} \right\} \Big|_{\theta_{\widehat{\mathbf{m}}} = \gamma_{\widehat{\mathbf{m}}}^{*}} = 0, \quad \widehat{\mathbf{m}} = 1, \dots \widehat{\mathbf{M}}$$
 (9.1)

where $\theta=(\underline{\theta}_{\widehat{\mathbf{m}}},\underline{\theta}^{+})$, in which $\theta_{\widehat{\mathbf{m}}}$ is the particular parameter to be estimated, with

$$\ell_{J-SCF}^{(10)} \equiv \log \left\{ \langle F_{J} \{ \underline{X} | \underline{S}(\theta_{\widehat{m}}, \underline{\theta}') \} \rangle_{\underline{\theta}'} / F_{J}(\underline{X} | 0) \right\}$$
(9.1a)

now the (logarithms of the) generalized likelihood ratio representing the test states $H_1(S\oplus N)$ vs. $H_0(N)$, which is the familiar "on-off" or "signal and noise" vs. "noise alone" situation in detection theory.

In a similar way we see that the rule (A.3-6) for obtaining the Bayes estimators for the QCF (A.3-5) can be equivalently expressed in terms of another likelihood ratio by

$$\{\gamma_{\mathbf{m}}^{\star}\}_{\mathbf{QCF}} = \{\int_{\Omega_{\widehat{\mathbf{m}}}} \theta_{\widehat{\mathbf{m}}} \sigma(\theta_{\widehat{\mathbf{m}}}) e^{2 \int_{\mathbf{M}}^{\mathbf{M}} (\mathbf{X} | \theta_{\widehat{\mathbf{m}}}) QCF} d\theta_{\widehat{\mathbf{m}}} \}, \qquad (9.2)$$

where now

$$\ell_{J-QCF}^{(21)} \equiv \log \langle F_{J}(\underline{X}|\underline{S}(\theta_{\widehat{m}},\underline{\theta'}))\rangle_{\underline{\theta'}} / \langle F_{J}(\underline{X}|\underline{S})\rangle_{\underline{\theta}}.$$
 (9.2a)

This likelihood ratio, $\ell_{J-OCF}^{(21)}$, represents the binary-signal detection situation where we test $H_2(S_{\widehat{m}} \oplus N)$ vs. $(H_1(S \oplus N), i.e., S^{(2)} \equiv \underline{S}_{\widehat{m}}, \text{ that portion of } \underline{S} \text{ containing } \theta_{\widehat{m}} \text{ only, and } \underline{S}^{(1)} = \underline{S}, \text{ cf. } (9.2a).$ Thus we note that γ_{m-SCF}^* is a linear <u>functional</u> of $\ell_{J}^{(10)}$, cf. (9.1), whereas the $\{\gamma_{m-QCF}^*\}$ are (monotonic) nonlinear functionals of $\ell_{J}^{(21)}$, cf. (9.2).

Having cast these estimators in likelihood ratio form, we can now proceed directly as in the previously developed optimum threshold detection theory ([13], [16], Appendix A.3; [15], Appendix) to obtain the needed approximate forms for $\ell_J = g_J^*(\underline{x}, \theta_{\widehat{m}}) + t_J(\underline{x}, \theta_{\widehat{m}}) = g_J^*$, where it is shown that as $J \to \infty$, $t_J \to 0$ prob. 1 on H_1 vs. H_0 (or H_2 vs. H_1), i.e., $\ell_J = g_J^*$ is AO as $J \to \infty$, cf. (7.12) et seq. Because of the functional relations (9.1), (9.2), χ^* is likewise then AO, as well as LOBE, i.e., is AOBE (asymptotically Bayes estimator). Now specifically (9.1), (9.2) become

$$\frac{\text{SCF:UMLE}}{\partial \theta_{\widehat{\mathbf{m}}}} \left[\frac{\partial \log \sigma_{\mathbf{1}}(\theta_{\widehat{\mathbf{m}}})}{\partial \theta_{\widehat{\mathbf{m}}}} + \frac{\partial g_{\mathbf{J}-SCF}^*}{\partial \theta_{\widehat{\mathbf{m}}}} \right]_{\theta_{\widehat{\mathbf{m}}} = \gamma_{\widehat{\mathbf{m}}}^*} = 0; \qquad (9.3)$$

$$\underline{QCF:LMSE} \quad \{\gamma_{\widehat{m}}^{*}\} = \{\int_{\Omega_{\widehat{m}}} \theta_{\widehat{m}} \sigma_{1}(\theta_{\widehat{m}}) e^{g_{\widehat{J}}^{*} - QCF} d\theta_{\widehat{m}} , \qquad (9.4)$$

where the g_J^* are given explicitly in the next section. A sufficient condition that g_J^* is AO is that g_J^* be asymptotically normal, with means ${}^{\mp}[var_0g_{J-SCF}^*,var_1g_{J-QCF}^*]$, ${}^{\mp}=H_0$, H_1 ; or H_1 , H_2 , and variances $var_0g_{J-SCF}^*,var_1g_{J-QCF}^*$, respectively. [Note, also, that because we have assumed the desired signal is a priori known to be [55] present, $log_{\mu}=log(p/q)$ of the detection forms, cf. (2.4) of [20], are omitted here in the formation of $\ell_J^{(10)}$, $\ell_J^{(21)}$ above.]

B. AO Forms of the gt

From Sec. 3.2 of [15], and from [20] suitably adapted, we have specifically the needed structures of g_J^* required in (9.3), (9.4) for determining the associated AO LOBE's. Specializing to the practically useful cases of independent data (i.e., noise) samples and locally stationary noise processes, which we shall henceforth assume in the remainder of our treatment, we have once more, cf. Sec. 7.2

(1) Coherent Reception (SCF)

$$g_{J-coh}^* = -\frac{1}{2} \sigma_{oJ-coh}^{*2} - \sum_{j=0}^{J} a_{oj} s_{j}$$
 (9.5)

(2) Incoherent Reception (SCF)

$$g_{J-inc}^{\star} = -\frac{1}{2} \sigma_{oJ-inc}^{\star 2} + \frac{1}{2!} \int_{ij}^{J(=MN)} (\ell_i \ell_j + \ell_i \delta_{ij}) \langle a_{oi} a_{oj} s_i s_j \rangle_{\underline{g}},$$
 (9.6)

(3) Composite Reception (SCF)

$$g_{J-comp}^{\star} = g_{J-coh}^{\star} + g_{J-inc}^{\star}, \qquad (9.7)$$

where, as in Sec. 7 et seq., $a_{0i} = a_{0}(t_{i})$, etc. is a (normalized) signal amplitude, $s_{i} = s(t_{i}, \theta, \underline{\theta}')$ is a normalized signal waveform, such that $\langle s^{2} \rangle = 1$, and $\ell_{i} = d \log w_{i}(x_{i})/dx_{i}$, $\ell_{i}' = d\ell_{i}/dx_{i}$, where $w_{1}(x_{i}) = first$ order pdf of the interference (nongauss + gauss) above, in the usual way. Specifically, here and henceforth in Sec. 9, $a_{0j} = a_{0n}^{(m)}$, $s_{j} = s_{n}^{(m)}$, since j = m, n, cf. Sec. 7.2 et seq. In this way we include the spatial, as well as temporal description of the sampled waves. Thus i = (m, n); j = (m', n'), when used in the incoherent cases.

The proper bias terms are given by $\hat{B}_{J}^{\star} = -\sigma_{oJ}^{(\star)^2}/2$, cf. (7.12), or (7.9), (7.10), with (7.11), etc. [We must be careful here and henceforth to distinguish between the various averaged signal parameters $(\underline{\theta}')$ and the unaveraged ones $\{\theta_{\widehat{m}}\}$, which are to be estimated.]

Similarly, for the QCF the associated g_3^* needed in (9.4) are

(1) Coherent Reception (QCF)

$$g_{J-coh}^{(21)*} = B_{J-coh}^{(21)*} - \sum_{j}^{J} \ell_{j} [\langle a_{oj} s_{j} \rangle_{\underline{\theta}}, -\langle a_{oj} s_{j} \rangle_{\underline{\theta}}] . \qquad (9.9)$$

(2) Incoherent Reception (QCF)

$$g_{J-inc}^{(21)*} = B_{J-inc}^{(21)*} + \frac{1}{2!} \sum_{ij}^{J} (\Delta \rho)_{ij} (\ell_i \ell_j + \ell_i^{\dagger} \delta_{ij}),$$
and (3)
$$g_{J-comp}^{(21)*}|_{QCF} = (g_{J-coh}^{(21)*} + g_{J-inc}^{(21)*})_{QCF},$$
(9.10)

cf. (9.7). Here $(\Delta \rho)_{ij} = \langle a_{0i}a_{0j}s_is_j \rangle_{\underline{\theta}}$, $-\langle a_{0i}a_{0j}s_is_j \rangle_{\underline{\theta}}$, and the proper biases for AO behavior are

$$B_{J-coh}^{(21)*} = -\frac{\sigma_{J-coh}^{(21)*}}{2} = -\frac{L^{(2)}}{2} \sum_{j}^{J} \left[\langle a_{oj} s_{j} \rangle_{\underline{\theta}}^{2}, -\langle a_{oj} s_{j} \rangle_{\underline{\theta}}^{2} \right]$$
(9.11)

$$B_{J-inc}^{(21)*} = -\frac{\sigma_{J-inc}^{(21)*}}{2} = -\frac{1}{8} \sum_{i,j}^{J} \{(L^{(4)}-2L^{(2)})^{2}\} \delta_{ij} + 2L^{(2)}\} \Delta R_{ij}, (9.12)$$

where
$$\Delta R_{ij} = \langle a_{0i}a_{0j}s_{i}s_{j}\rangle_{\underline{\theta}}^{2}$$
, $-\langle a_{0i}a_{0j}s_{i}s_{j}\rangle_{\underline{\theta}}^{2}$ ($\neq (\Delta \rho)_{ij}$).

C. Suboptimum Estimators

We parallel the treatment in [15], [20], where g_J^* is replaced by the suboptimum g_J , now appropriately adapted to the estimation structures (i.e., hypothesis test) required by the SCF, QCF, etc. We illustrate the procedure with the case of SCF and the use of simple (cross- and auto-) correlation receivers, well-known to be threshold optimum in gauss interference. Now (9.5)-(9.7) reduce to this case directly on setting $\ell_i = -x_i$ ($\ell_i = -1$), and $\ell_i = -x_i$ (2) $\ell_i = -1$), and $\ell_i = -x_i$ (3), we make the same substitutions, in both the bias and data portions of g_{OCF}^* .

In the important cases of clipper-correlators, which are optimum in Laplace noise and which are well known to be effective against non-gaussian noise, cf. [15], we replace ℓ_i by sgn \mathbf{x}_i (and omit ℓ_i), with appropriate modifications of the bias terms, cf. Sec. 3.3 of [15] for details. The basic pdf's for which these various correlation detectorforms (and hence estimator forms via (9.1), (9.3)) are optimum, are respectively given by $\mathbf{w}_1(\mathbf{x})$ in Sec. 7.5 above.

Finally, it is important to note that these LOBE's [obtained from (9.1), (9.3)] are only practically AO and LOB when $\theta^2(=< a_0^2>)$ is sufficiently small, as in detection, i.e., there exists an $< a_0^2>_{max}$ (<< 1), with $< a_0^2> << a_0^2>_{max}$, such that for $\overline{a_0^2}> < a_0^2>_{max}$ the threshold optimal character of g_0^* breaks down and becomes suboptimum. At larger input signals g_0^* ($\neq g_1$ now) and hence $\gamma_m^* + \gamma_{\widehat{m}}$, etc., may or may not be monotonically better, in absolute terms, than g_0^* , $\overline{a_0^2}<\overline{a_0^2}$ and $\overline{a_0^2}$. Usually, for sufficiently strong signals, g_0^* (now g_1) is absolutely better, unless the information-bearing portion of the signal is destroyed by the algorithm itself, e.g., clipping

destroys waveform detection/estimation, but not phase parameter extraction, for example.

9.2 An Example: Amplitude Estimation

with the SCF: Coherent Reception

The simplest useful example is that of estimating the scale (or amplitude) $\sim a_0$ of an otherwise fully known signal waveform received in (generally) nongaussian interference. This is also the problem of <u>estimating</u> signal intensity $\hat{I}_{\infty}^* = \hat{a}_0^{*2}$.

Here we have $\theta_{\widehat{m}} = \theta_1 = a_0$, with $\underline{\theta}'$ all other relevant signal parameters. With coherent reception, signal epoch ε is precisely known, e.g., $\varepsilon = \varepsilon_0$, and we select $\langle s_j \rangle_{\varepsilon} = s_{j-max} = \sqrt{2}$, with independent noise samples. Thus, we write $\langle s_j \rangle_{\underline{\theta}'} = s_{max} = \sqrt{2}$ here. Physically, our present example can represent slow fading, whose changes are negligible over the data acquisition period.

Accordingly, (9.5) reduces directly to

$$(g_{J-coh}^*)_{SCF} = -L^{(2)} a_{oJ}^2 - a_{o} \sqrt{2} \sum_{j=1}^{J} \ell(x_i).$$
 (9.12)

Applying this to (9.3), with $\theta_{\rm m}$ = $\theta_{\rm 1}$ = $a_{\rm o}$, gives

$$\left[\frac{\sigma_{1}(a_{0})'}{\sigma_{1}(a_{0})} - L^{(2)}a_{0J}^{2} - a_{0}\sqrt{2}\sum_{j} \ell(x_{j})\right]_{a_{0} = \hat{a}_{0}^{*}} = 0, \qquad (9.13)$$

whose solution yields the desired estimator $\hat{a}_{0}^{*}(x)$.

To proceed further, we need to know the a priori pdf of $\theta_1(=a_0)$, e.g., $\sigma_1(a_0)$. Let us assume, then, that $\sigma_1(a_0) = 1/\Delta a_0$, a uniform pdf on $(0-\le a_0\le a_0+)$, and zero elsewhere. Consequently, $\sigma_1'(a_0)=0$, $0\le a_0\le \Delta a_0$ (the contribution of the δ -functions at 0, Δa_0 provide no meaningful solutions). Solving (9.13) gives

$$\begin{vmatrix} \hat{a}_{0}^{*}(\underline{x}) \\ \text{coh} \\ \text{SCF} \\ \text{uniform} \end{vmatrix} = -\sum_{j}^{J} \ell_{j}/J\sqrt{2}L^{(2)}$$
, $\overline{a_{0}^{2}} < 1$ (9.14)

The optimum threshold estimate of signal intensity is accordingly

$$\hat{I}_{S}^{*} = (\hat{a}_{O}^{*})^{2} = (\sum_{j} \ell_{j})^{2}/2J^{2}L^{(2)}^{2}.$$
(9.15)

Here \hat{a}_0^* , and \hat{a}_0^{*2} are UMLE for the assumed uniform pdf of a_0^* , and simultaneously CMLE for any (meaningful) pdf of a_0^* . Moreover, $\langle a_0^* \rangle_1 = \overline{a_0}$, because $\langle l \rangle_{H_1} = \langle l \rangle_1 = -\sqrt{2} |a_0^*| = -\sqrt{2} |a_0^*|$, and we can show that

$$\operatorname{var}_{1} \hat{a}_{0}^{*} \doteq \overline{a_{0}^{2}} (1 + O(\frac{1}{J})) - \overline{a_{0}^{2}} \doteq \operatorname{var} a_{0}|_{J^{>>}1}, \quad (\overline{a_{0}^{2}} << 1)$$
 (9.16)

from $\langle \ell_i \ell_j \rangle_1 = 2\overline{a_0}^2 L^{(2)^2}$ ($i \neq j$); $= \ell^2 + \overline{a_0}^2 L^{(2,2)}$ ($i \neq j$). Thus, \hat{a}_0^* is unconditionally unbiased, e.g., $\langle \hat{a}_0^* \rangle_1 = \overline{a_0} (= \Delta a_0/2)$ and $\lim_{J \to \infty} \hat{a}_0^* \to a_0$, from (9.16).

The "smallness" condition on $\overline{a_0^2}$ is (from Sec. 6.4, Eq. (6.71) of [15])

$$\overline{a_0^2} << \frac{1}{2} \frac{var_0 \ell}{var_0 \ell^2} = L^{(2)}/[L^{(2,2)} - 2L^{(2)^2}] = x_0; \quad L^{(2,2)} = <(w_1'/w_1)^4 >$$
(9.17)

Here, from $\sigma_1(a_0) = 1/\Delta a_0$ we have $\overline{a_0^2} = (\Delta a_0)^2/3$ in (9.16) to establish an upper bound $(\overline{a_0^2})_{max}(<<1)$ for which $\overline{a_0^2} \leqslant (a_0^2)_{max}$ and $\hat{a}_0^*(\underline{x})$ is then LOBE and AO.

As a numerical example, consider a Class A noise for which $\Gamma_A^1 = 10^{-4}$, $A_A = 0.5$, so that $x_0 = 1.7 \cdot 10^{-4}$ (from Fig. 5.6 of [15]): $x_0 = 0.5$ and $x_0 = 0.5$ with the result that $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ are $x_0 = 0.5$ are $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$ are $x_0 = 0.5$ are $x_0 = 0.5$ and $x_0 = 0.5$

9.3 Discussion

First, let us note that our illustrative example is readily extended to the estimation of waveform itself $S = \{S_j\}: \theta_{\widehat{m}} \to S_{\widehat{m}}$, and we get a series of relations like (9.13), if $\sigma(\underline{\theta}) = \widehat{m-1}\sigma_1(\theta_{\widehat{m}})$; otherwise one has M-coupled equations to be solved for $\theta_{\widehat{m}} = S_{\widehat{m}}$. Also, if the QCF is used, the structure

of the optimal estimator is generally much more complex, due to the non-linear functional relation between g_j^* and γ_j^* , cf. (9.2). If the parameters to be estimated appear functionally in the waveform $S(t,\underline{\theta})$ then one obtains functional solutions for the estimators.

The important general result here (for all signal levels) is the explicit functional relation between the (optimum) detection algorithms and the resultant optimum estimators, cf. (9.1), (9.2): detector structure provides the initial analytic relation, from which the estimator is the appropriate functional relation, the form of which depends on the choice of cost function. Thus, to carry out the derivation of the desired estimator, we must start with the appropriate detector algorithm [16]. The evaluation of estimator performance is provided by using the estimator, $\gamma_{\rm J}^{(\star)}$, in the expressions for the average error (or risk), cf. (A.3-2).

Finally, one important extension of the analysis outlined above is to the frequently occurring situation (mentioned at the beginning of Sec. 9) necessitating joint detection and estimation, when it is not known precisely that the desired signal is present [59].

10. Concluding Remarks

In the preceding section we have outlined a comprehensive methodology for handling the general problems of acoustic signal transmission and reception in and through complex underwater media. Our approach has included the following important channel features: (1) inhomogeneous volumes and interfaces, both deterministic and random; (2) spatial as well as temporal effects; (3) the inherently nongaussian character of major components of the random fields arising both from ambient sources and produced by a variety of scattering mechanisms; (4) space-time processing by the distributed arrays which couple source and receiver to the medium; and (5) threshold detection and estimation of signals in these generally nongaussian environments.

In addition to the above topics we have touched upon a variety of specific methods and techniques for channel modeling, including operator formalisms and diagrammatic methods. A number of new results are also discussed: poisson field models and statistics, canonical threshold

estimators and estimates, for example, as well as measures of threshold detector performance which specifically incorporate the rôles of the receiving array, signal structure (with doppler smearing and fading), and mixed gaussian and nongaussian noise components in the interference. Spatial processing gains are seen to be achieveable in the usually encountered nonuniform noise fields by suitable spatial sampling, for both detection and estimation, in addition to the capability of beam formation. Suboptimum, as well as optimum processing algorithms are noted.

The preceding sections are intended to be a general guide to the formulation and treatment of specific propagation and signal processing problems, wherein attention is called to the many physical features which may have to be taken into account in individual problems. The accompanying references should provide the in-depth analysis and/or specific results needed at each stage of the problem in question. As stated at the beginning, our general aim here has been to provide an overview, keyed to the realities of specific applications, and at the same time to exhibit a unifying and interdisciplinary approach, within which the many special modeling and processing tasks of underwater acoustic interests are embedded.

Appendix A.1 Diagram Representations and Solutions [2]

Diagram representations and methods are very useful and compact forms for describing field interactions as is well known [23], [41]. For some of our subsequent results we shall include diagram equivalents, along with the corresponding analytic expressions, defining the symbol "vocabulary" in a consistent way. For our ensemble results [(3.4)-(3.9)], we write the equivalent ensemble diagrams:

$$FOR: -G_{T} \xrightarrow{\widehat{Q}} (\{\alpha\} \equiv \underbrace{T} \xrightarrow{\widehat{M}} \underbrace{\{\alpha\}}_{\{\widehat{Q}\}}) : (A.1a)$$

$$\frac{\text{FD}}{\{\alpha\}} = \frac{1}{\{\alpha\}} + \frac{1}{\{\alpha\}} \circ \text{(Eq. (3.4);} \quad (A.1b)$$

$$\frac{\text{FOS}:}{\begin{cases} \vdots \\ \text{Eq. (3.8)} \end{cases}} = \frac{\hat{1}}{\hat{1} - \cancel{\longrightarrow} \circ} \xrightarrow{\text{H}} ; \xrightarrow{\text{H}} = \cancel{\longrightarrow} (\hat{1}) = -\hat{M}_{\infty}(G_{\uparrow}) \text{ (A.1c)}$$

PTSS:
$$= \hat{1} + \longrightarrow 0 + \longrightarrow 0 + \dots]$$
 H (A.1d)

where we define the symbols as we go along. Thus, \longrightarrow ($\equiv M_{\infty}$) denotes the "feedforward" operator, while \longleftarrow o($\equiv \{Q\}$) denotes the (ensemble) "feedback" or scattering operator. The ensemble Feynman diagram (FD) equivalent of the FOR (A.1a) is just (A.1b), with (A.1c) giving the corresponding ensemble FOS, and (A.1c) the PTSS, obtained by iteration of (A.1b) or from (3.9) directly.

A.1-1: Stochastic Solutions

As we have noted above [cf. Eq. (2.8) et seq.] the solutions to the Langevin equation (2.8) are the various statistics of the field ι , e.g., here the moments $\langle \iota \rangle$, $\langle \iota_1 \iota_2 \rangle$, etc.

Let us consider first the mean field $<\alpha>$. Taking the average of (3.9), for example, gives directly

$$\langle \alpha \rangle = \alpha_{\text{H}} + \sum_{n=1}^{\infty} \langle \hat{\mathbf{\eta}}^{(n)} \rangle \alpha_{\text{H}},$$
 (A.2)

which shows that all $(n\frac{th}{-})$ order moments of \hat{Q} are required. The diagram equivalent of (A.2) is, from (A.1d),

$$\left\{ \begin{array}{l} \Longleftrightarrow = \left[\hat{1} + \overrightarrow{\hat{M}} \right] + \overrightarrow{\hat{M}}_{1} < \widehat{Q}_{1} + \cdots \right] & \text{ with } \\ \bullet = < \mathbb{Q}; & \bullet & \bullet \\ \end{array} \right.$$

$$\left\{ \begin{array}{l} \Longleftrightarrow \\ \alpha_{H} \end{array} \right.$$

$$\left(\begin{array}{l} \Longleftrightarrow \\ (n) \end{array} \right) = < \widehat{Q}_{1} \cdots \widehat{Q}_{n} > .$$

$$\left(\begin{array}{l} \Longleftrightarrow \\ (A.2a) \end{array} \right)$$

There is a second approach which leads to an equivalent deterministic formulation, which is usually more effective for approximations. Defining a deterministic operator $\hat{Q}_1^{(d)}$ by

$$\hat{Q}_1^{(d)} < \infty \equiv < \hat{Q}_{\infty}$$
, (A.3)

we may again average (3.4) and use (A.3), to get

$$< \infty = \alpha_{H} + \hat{M}\hat{Q}_{1}^{(d)} < \infty = (\hat{1} - \hat{\eta}^{(d)})^{-1} \alpha_{H} = \alpha_{H} + \sum_{n=1}^{\infty} (\hat{\eta}^{(d)})^{(n)} \alpha_{H}, \quad (A.4)$$

or, in diagram form,

$$= \frac{1}{\hat{q}_{1}} + \frac{d}{\hat{q}_{1}} = \frac{\hat{1}}{\hat{1} - \frac{d}{d}} + \frac{1}{\hat{q}_{1}} + \frac{1}{\hat{q}_$$

respectively.

Equation (A.4) is a form of <u>Dyson's equation</u> (DE), where now the equivalent deterministic scattering operator, $\hat{Q}^{(d)}$, or (EDSO), is the analogue of the "mass-operator" (\hat{Q}) in quantum electrodynamics (cf. ¶60,

[24]; [23]). The EDSO is always an <u>integral</u>, or global <u>operator</u>, over all pertinent space and time, and has the form

$$\hat{Q}_{1}^{(d)} = \sum_{m=0}^{\infty} \hat{A}_{m}(\underline{R}, t | \underline{R}', t') : \hat{A}_{0} = \langle \hat{Q} \rangle ; \hat{A}_{m} = \langle \hat{Q} \hat{M} \hat{Q}^{(m)} \rangle - \sum_{j=1}^{m \ge 1} \hat{A}_{m-j} \langle (\hat{M} \hat{Q})^{(j)} \rangle,$$
(A.5)

which at once permits a hierarchy of approximate forms, by stopping with term $(m \ge 1)$. The diagram equivalent of (A.5) is

$$\hat{Q}^{(d)} \equiv \longleftarrow = [\bullet]_{m=0}^{+} [\longleftarrow - \longleftarrow]_{m=1}^{+} [\longleftarrow - \longleftarrow + \longleftarrow]_{m=2}^{+} \cdots$$
(A.5a)

Similar concepts may be applied to higher-order moments of the field. In fact, we may write the following FOS for the second-order, second moment of the field:

$$\langle \alpha_{1}^{\alpha} \alpha_{2} \rangle = (\hat{1} - \hat{\eta}_{12}^{(d)})^{-1} (\hat{1} - \hat{\eta}_{1}^{(d)} \hat{\eta}_{2}^{(d)}) \langle \alpha_{1} \rangle \langle \alpha_{2} \rangle; \hat{\eta}_{12}^{(d)} \equiv \hat{M}_{1} \hat{M}_{2} \hat{Q}_{12}^{(d)},$$
(A.6)

where the EDSO, $\hat{Q}_{12}^{(d)}$, is defined by

$$\hat{Q}_{12}^{(d)} < \alpha_1 \alpha_2 > \equiv < \hat{Q}_1 \hat{Q}_2 \alpha_1 \alpha_2 > , \qquad (A.6a)$$

cf. $\hat{Q}_1^{(d)}$, (A.3) above. Equation (A.6) is a second-order Dyson equation, analogous to a form of the Bethe-Salpeter equation in quantum electrodynamics ([21]; ¶60, [24]). Analogous to (A.4a), we have

$$\begin{cases} \Leftrightarrow = & \text{ with } + & \text{if } & \text{if }$$

with obvious extensions of the second and third diagrams in (A.4a). The EDSO, (A.6a), is the extension of (A.5a):

$$\hat{Q}_{12}^{(d)} = \underbrace{\overset{d}{\downarrow}}_{12} = \begin{bmatrix} & & & & \\ & 1 & 2 \end{bmatrix}_{12} - \underbrace{\overset{d}{\downarrow}}_{12} + \underbrace{\overset{d}{\downarrow}}_{12} + \underbrace{\overset{d}{\downarrow}}_{12} \end{bmatrix}_{m=0} + \dots$$
 (A.7b)

Higher order Dyson forms of the FOS, like (A.6), can be constructed for $<\alpha_1...\alpha_m>$. In addition, various approximations can be instituted, in $\hat{\mathbb{Q}}^{(d)}$ (series modification) where the series are infinite (for some form of strong scattering), or by <u>truncation</u>, for "weak" scattering (1st-Born approximation, for example [cf. (5.1)]), where one stops at the dotted line in (A.1d), (A.2a), with n=1 in (A.4a), etc. Unfortunately, space does not permit us to pursue these topics further here.

A.1-2. Received Waveforms

The expression for the received waveforms, after sensing by the receiving array (\hat{R}) may be similarly diagrammed. We start with (A.1), etc. applied to (4.1) and (4.2). Thus, we have

$$\hat{\ell} = \frac{}{\hat{1} - } \xrightarrow{\circ} ; \langle \hat{\ell} \rangle = \frac{}{\hat{1} - } \xrightarrow{\circ} , \text{ etc. cf. (A.4)},$$

(A.8a) and, in PTSS form, cf. (A.1d), Eq. (4.2):

$$\hat{K}_{g} = \begin{bmatrix} & & & & & & \\ & 11 & 22 & & & \\ & & 11 & 22 & & \\ & & & & 11 & 22 & & \\ & & & & & & \\ & & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & \\ & & & \\ & & \\ & & & \\$$

The diagram of the mean field is given at once by (A.2a), or (A.4a), on multiplying the receiver operator \hat{R} since $\langle X \rangle \approx \hat{R} \langle D \rangle$. The diagram for (4.5), in PTSS form for $\langle X_1 X_2 \rangle$ is, similarly by direct expansion and averaging,

Eq. (4.5):

$$\langle x_1 x_2 \rangle = \hat{R}_1 \hat{R}_2 [\xrightarrow{1}_2 + \{ \rightarrow (\rightarrow + | \rightarrow + \dots)_2 + \xrightarrow{2}_2 (\rightarrow + | \rightarrow + \dots)_1 \} + \dots \}]$$

$$+ \{ \xrightarrow{11}_{22} + | \xrightarrow{1}_{22} + 2 \xrightarrow{2}_{21} \xrightarrow{1}_{11} + \dots \}]$$

$$(A.8c)$$

from which the various higher-order moments of $\hat{\mathbb{Q}}$ (at points 1,2) appear clearly.

Appendix A.2 Canonical Scattering Operator Formalisms: Basic Decomposition Principle

For this it is necessary now to use a canonical formulation of the scattering operator, \hat{Q} , and in particular, the author's concept of a <u>Basic Decomposition Principle (BDP)</u>, which permits both a statistical and physical quantification of this operator in terms of differential scattering elements (dse's) which, in turn, can be identified with specific physical mechanisms. Again, space limitations allow only the briefest summary of results below.

We write the following canonical expressions

$$\hat{Q} + \hat{Q}_{M}(\underline{r}, t | \underline{r}', t') = \int_{-\infty}^{\infty} dt' \int_{Z_{R}} dN(\underline{Z} | \underline{r}', t'; t)_{M} h_{M}(t - t', t | \underline{r}', \underline{r}, ...)$$

$$\hat{d}_{om}(\underline{r}, t | \underline{r}', t')(\underline{r}', t'), \qquad (A.2-1)$$

where

 dN_{M} = the number of <u>available</u> dse's in the domain dA_{M} ; fdN() is a "counting" functional.

 \hat{d}_{ON} = the <u>local interaction operator</u>, determined by the Langevin equation (and boundaries), relating the dse and incident field.

(A.2-la)

Here \underline{Z}_R is the space of all radiation "events," e.g., scattering, and \underline{Z} (= $\underline{Z}_R \times \underline{Z}_{S \times \Theta}$) is the total random variable and parameter space ($\underline{\Theta}$); ($\underline{r},\underline{r}'$) are vectors in the coördinate system of the scattering elements, Λ_M . In most cases we use the <u>basic point-scatter model</u> (BPSM) for h_M :

$$h_{M}(t',t|\underline{r}',...) = \gamma_{o}(\underline{r}',t-t')_{M}\delta(t-\Delta t_{dS}); \quad \Delta t_{dS} = f(t',t;\underline{r}',\underline{r}), \quad (A.2-2)$$

where $\Delta t_{dS} = t_{dS}$ -a is the displacement (in time units) due to doppler of the dse, away from some fixeJ position in Λ_{M} and γ_{O} (> 0) is a "cross section" which depends on the nature of the medium (or boundaries).

The inhomogeneous character of the medium (and boundaries) is of two types, A, B. We have, with defining conditions:

A. The Inhomogeneum:
$$(0 \neq) \lambda_0 |\nabla \mu| + \frac{\lambda_0}{c_0} \left| \frac{\partial \mu}{\partial t} \right| < O(1), c_0 = \lambda_0 f_0,$$
 (A.2-3a)

B. The Discontinuum:
$$\lambda_0 |\nabla \mu| + \frac{1}{f_0} \left| \frac{\partial \mu}{\partial t} \right| >> O(1), \quad \mu(\underline{r},t) = \varepsilon(\underline{r},t), \text{ etc. } (A.2-3b)$$

where the latter represents a "hard" boundary and the former represents comparatively small changes in the continuous medium. The Discontinuum can be localized, e.g., particles, bubbles, etc., or distributed, like the air/water interface. It is then seen that

A:
$$\hat{d}_{OM}|_{Inhomog} = \delta(\underline{r}-\underline{r}')\hat{Q}_{coeff}.(\underline{r}',t');$$
 B: $\hat{d}_{OM}|_{Discont}.=\delta(\underline{r}-\underline{r}')\hat{1},$ (A.2-4)

where \hat{Q}_{coeff} is the coefficient of $\mu(r,t)$ in the Langevin equation, e.g., of $\epsilon(r,t)$ in (3.2), viz., $c_0^{-2}(\partial^2/\partial t^2)$, etc. In consequence of the assumed lack of motion of the fluid as a whole, there is no doppler associated with $\mu(r,t)$, whereas there is always a (random) doppler associated with the Discontinuum elements, about some equilibrium position of the ocean wave surface, or particle displacement in the volume. (For a general treatment of these important doppler effects, see [36].)

Finally, to anatomize the various types of possible scattering interaction, we invoke the following Basic Decomposition Principle (BDP) for the number of available dse's, dN, [1], [2]. The BDP states that:

"From the viewpoint of radiative interactions among the dse's which constitute the medium, the argument (dN, cf. (A.2-1)), of the "counting" functional ($\int_{ZR} dN() = N$, (A.2-1a)), can be decomposed into $k = 1, 2, ..., \infty$ distinct subsets of available dse's for radiation, each subset of which contains respectively distinct and independent k-couple radiating elements only." (A.2-5)

This may be expressed, after some manipulation, in terms of (radiative) mean values, e.g., $<>_R$, and fluctuations of the various <u>densities</u>, $v^{(k)}$, of different classes $(k \ge 0)$ of dse's, by

$$\frac{\text{BDP}:}{\langle v(\underline{r},t) \rangle} = \sum_{k=1}^{\infty} v^{(k)}(\underline{r},t) = \langle v(\underline{r},t)^{(0)} \rangle_{R} + \sum_{k=1}^{\infty} \Delta v^{(k)}(\underline{r},t);$$

$$\langle v^{(0)} \rangle_{R} = \sum_{1}^{\infty} \langle v^{(k)} \rangle_{R}, \quad \Delta v^{(k)} = v^{(k)} - \langle v^{(k)} \rangle_{R}, \quad (A.2-6)$$

since $\hat{N}\delta = v(\underline{r},t)$, $\underline{r} \in \Lambda_M$, when M=S or V (or B), of course. Here (k=0) represents the average number of dse's (discrete or continuous) engaged in (re-radiation); k=1 denotes the aggregate of independent, uncoupled dse's, while $k \ge 2$ are the sets of pairs (k=2), triples (k=3), etc. of dse's available for (multiple) scattering. It follows directly from the BDP that all sets $(k \ge 0)$ are statistically independent. From this it can be shown, furthermore, cf. [21], that the fluctuations $\Delta v^{(k)}$ obey (zero-mean), $k^{\underline{th}}$ order poisson statistics, under very general conditions, reducing to gaussian statistics whenever the process densities are sufficiently great. Figure A.2-1 gives a schematic illustration of the decomposition principle presented here.

A.2-1. Operator Structure and Radiation Event Statistics

Further insight into the nature of the BDP, (A.2-5), and the canonical structure of the inhomogeneity operator \hat{Q}_M , (A.2-1), may be gained from the following. It can be shown for small regions $d\Lambda_M$ in which dQ_M is accordingly the scattering operator, that the BDP can be anatomized as follows, cf. Fig. A.2-1:

$$(\underline{BDP}): \qquad d\hat{Q}_{M} = \sum_{k=1}^{\infty} d\hat{Q}_{M}^{(k)} = \hat{\mathcal{L}}_{dA} \left\{ \sum_{k=1}^{\infty} a^{(k)} \gamma_{OM}^{(k)} Q^{(k)} \right\}, \tag{A.2-7a}$$

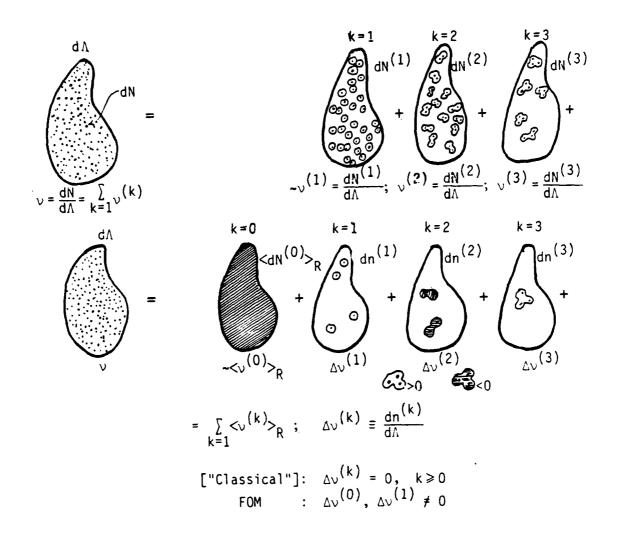


Figure A.2-1 Decomposition Principle: $dv = \sum_{k=1}^{\infty} v^{(k)}$; $dN = \sum_{k=1}^{\infty} dN^{(k)}$: scattering domain resolved into hierarchy of independent, k-coupled scattering elements.

where $\hat{oldsymbol{\mathcal{X}}}_{d\Lambda}$ is a linear (operator) functional, with

$$(\underline{BDP}): \qquad \hat{Q}_{M} = \mathcal{L}_{\Lambda_{M}} \{ \sum_{k=1}^{\infty} a^{(k)}, (k) \vee (k) \}$$
(A.2-7b)

over the entire region (A_M) of scatterers. Here $y^{(k)} = y^{(k)}(\underline{r},t)$ is the density of k-coupled scattering elements ("quasi-particles") in A_M . The quantities $y^{(k)}_{OM}$, $a^{(k)}$ are respectively the scattering "cross-section" of

the $k\frac{th}{}$ - order "quasi-particles," and an appropriate weighting function, noted below. Thus, the BDP shows how a scattering region is decomposable into <u>independent</u> groups of interacting elements. [This decomposition has certain analogies with "Cluster Theory" in the statistical mechanics of a classical nonideal gas [42], but in any case we are not concerned here with the details of the k-tuple interactions.] Each k-tuple is entirely equivalent to a $k\frac{th}{}$ - order (classical) "quasi-particle," in the language of particle physics ([41], Chapter 2). Thus, (A.2-6) can be interpreted as a sum of different densities of "quasi-particles," which interact with the incident field to produce corresponding orders of single- and multiple-scatter.

Comparing (A.2-7b) and (A.2-1) specifies $\hat{\boldsymbol{\chi}}_{\Lambda_{\mathbf{M}}}$, namely,

$$\hat{Q}_{M} = \int_{-\infty}^{\infty} dt' \int_{\Lambda_{M}} \sum_{k=1}^{\infty} \sqrt{(k)} \gamma_{0}^{(k)} h_{F}^{(k)}(t-t',t|\underline{r}',\underline{r},...) \hat{d}_{0}(\underline{r},t|\underline{r}',t') \left[\underline{J}_{\underline{r}',t'}d\Lambda_{M} \right] (A.2-8)$$

where $a^{(k)} = h_F^{(k)}$ is a time-varying weighting function representing the reradiation process from the $k^{\underline{th}}$ -order "quasi-particles" (or scattering elements*). Again, \hat{d}_0 is a <u>local interaction operator</u> (A.2-la), determined by the particular Langevin equation (3.2), for example (A.2-4) et seq. The key result of the BDP is that it removes our having to specify the details of the radiation interactions <u>within</u> the $k^{\underline{th}}$ -order "quasi-particle." Instead, we have a linear combination of quantities $(-\gamma_0^{(k)})$, cf. (A.2-8), which are to be determined explicitly from the appropriate physics of the scattering situation in question.

At this point we consider further the <u>concept of "radiation events"</u> (\underline{Z}_R) (and their statistics), i.e., the interactions of the incident field with a quasi-particle, followed by emission, for given space-time

^{*}The vectors \underline{r} , \underline{r}' are defined in a coördinate system based in Λ_M , which is generally different from the coördinate system relating source (T), the "ideal" observer in the field (at $P(\underline{R},t)$), or the receiver (R), cf. Fig. A.2-5.

coördinates (S = (r,t)), i.e., $Z_{S\times\Theta}$ fixed. The classical description of scattering (Z_{class}) is thus analyzed into a two-component process: $Z_{class} = Z_{S\times\Theta} \times Z_R$, and the classical statistical averages are $Z_{class} = Z_{class} \times Z_{class} \times Z_{class}$, where $Z_{class} = Z_{class} \times Z_{class} \times Z_{class}$, where $Z_{class} = Z_{class} \times Z_{class} \times Z_{class} \times Z_{class}$, where $Z_{class} = Z_{class} \times Z_$

The importance of this decomposition is that it permits us to determine the overall statistics of the $v^{(k)}$. This is accomplished by our first noting that the <u>fluctuation</u> $\Delta v^{(k)}$ in the number of $k^{\underline{th}}$ - order radiation interactions in $\Lambda_{\underline{M}}$ is given by (A.2-6).

Furthermore, we have the identity

$$\sum_{k=1}^{\infty} \gamma_{0}^{(k)} v^{(k)} h_{F}^{(k)} = \gamma_{0}^{(0)} h_{F}^{(0)} \langle v^{(0)} \rangle_{R} + \sum_{k=1}^{\infty} \gamma_{0}^{(k)} h_{F}^{(k)} \Delta v^{(k)} = \Gamma^{(0)} + \sum_{k=1}^{\infty} \gamma_{0}^{(k)} h_{F}^{(k)} \Delta v^{(k)},$$
(A.2-9)

so that by (A.2-7b)

$$\hat{Q}_{M} = \hat{\mathcal{X}}_{\Lambda_{M}} (\Gamma^{(0)} + \sum_{k=1}^{\infty} \gamma_{0}^{(k)} h_{F}^{(k)} \Delta v^{(k)}). \tag{A.2-9a}$$

It can be readily shown [11] that the <u>radiation-event statistics</u> (RES) of the $\Delta v^{(k)}$ are zero-mean poisson, with the $n^{\frac{th}{h}}$ -order characteristic function (c.f.)

$$F_{n}^{(k)}(i\underline{\xi},i\underline{t}|Z_{S\times\underline{\theta}}|\Delta v^{(k)}) = \exp\left[\int_{\Lambda_{1...n}} R_{n}^{(k)}[\exp(i\sum_{\ell=1}^{n}\xi_{\ell}) - i\sum_{\ell=1}^{n}\xi_{\ell}]d\Lambda_{1...n}\right],$$

$$(\text{each } k \ge 1, \text{ any } n \ge 1); \quad \Lambda_{1...n} = \Lambda_{1} \cdot \Lambda_{2} \cdot ... \cdot \Lambda_{n}, \quad (A.2-10)$$

in which $i\underline{\xi}=i\xi_1,\ldots,i\xi_n$; $i\underline{t}=it_1,\ldots,it_n$ and $R_n^{(k)}=R_n^{(k)}(\underline{\lambda}_1,t_1;\ldots,\underline{\lambda}_n,t_n)\geqslant 0$, $(k,n\geqslant 1)$ is the associated process density. Since the radiation events for different k are independent, the complete (conditional) c.f. of the $\Delta v^{(k)}$, $k\geqslant 1$ is

$$F_n^{(\infty)} = \prod_{k=1}^{\infty} F_n^{(k)}. \tag{A.2-11}$$

The overall statistics of $\gamma^{(k)}_{F}h^{(k)}_{F}\Delta\nu^{(k)}$ are found from the $(n^{\frac{th}{t}}-\text{ order})$ Fourier transform of) $\langle F_{n}^{(k)}\rangle_{Z_{S(r)}}$, and for $\sum\limits_{1}^{\infty}\gamma^{(k)}_{O}h^{(k)}_{F}\Delta\nu^{(k)}$ from the F.T.

 $\{<F_n^{(k)}>_{Z_{S_n^{(k)}}}\}$. Thus, a central problem in applying our canonical development here is to^{θ} determine the process densities $R_n^{(k)}$ for the specific physical problem in question. We note some examples below is Sec. A.2-2.

A.2-2 Examples: FOM and Classical Scatter Models

We begin by considering

A. FOM Models:

As noted in Section 5, FOM models [28]-[30] are quasi-phenomenological in that:

- (i). they do not explicitly include boundary conditions, but rather subsume them in the $\Gamma^{(0)}$ and $\gamma_0^{(k)}h_F^{(k)}$ factors of the scattering operator, $\hat{\mathbb{Q}}_{\mathbf{M}}$, cf. (A.2-9);
- (ii). they permit only k = 1 (and hence k = 0) types of radiation interactions, namely, the so-called independent "point"-scattering. This latter constraint ensures only "weak-scattering" in volumes (M = V), and correspondingly for surfaces (S,B), e.g.,

$$\alpha = \alpha_{H} + \hat{M}_{\infty}\hat{Q}_{M}\alpha_{H}$$

in (3.9), cf. (5.1) also, with the added condition that all k-order ($k \ge 2$) interactions in \hat{Q}_M are omitted: no multiple scatter, cf. (5.8).

Thus, in the FOM cases we have the following relation for the governing process density $R_n^{(1)}$:

$$R_n^{(k)} \rightarrow R_n^{(1)} = \sigma_M^{(1)}(\underline{r}) \prod_{\ell=2}^n \delta(\underline{r}_{\ell} - \underline{r}_{\ell-1}), \qquad (A.2-12)$$

where now $\sigma_{M}^{(1)}$ = average density of scatterers in the illuminated region Λ_{M} , (M = S,B,V). From (5.9), (5.9a) we see that $\langle v^{(1)} \rangle = \sigma^{(1)}$, $v^{(k)} = 0$, $k \ge 2$.

The original FOM models for surface scatter [28]-[31] postulated a flat surface (i=0), but employed a dopplerized path delay, t_d , in $h_F^{(k)}$ which implicitly contained the effects of the moving surface. FOM model extensions, FOM^(1a), explicitly include the surface elevation i in the doppler delay [36], viz.

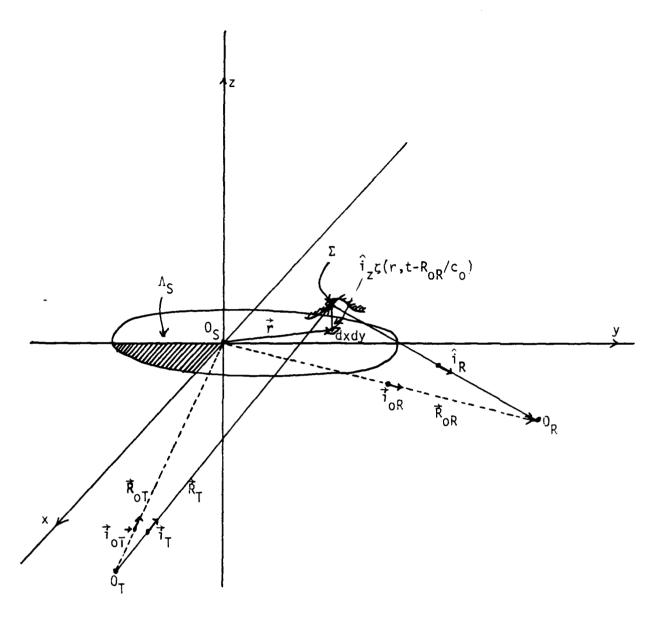


Figure A.2-2 Geometry of bistatic surface scatter.

$$t_{d} = T_{o} + (\hat{\underline{i}}_{T} - \hat{\underline{i}}_{R}) \cdot [\underline{r} + \underline{\zeta}(\underline{r}, t - R_{oR}/c_{o})]/c_{o}, \quad T_{o} = (R_{oR} + R_{oT})/c_{o};$$

$$\underline{\zeta} = \hat{\underline{i}}_{z}\zeta. \quad (A.2-13)$$

where $\underline{r} = \hat{\underline{i}}_x x + \hat{\underline{i}}_y y$ on Λ_S ($\zeta = 0$) and $\hat{\underline{i}}_T = \underline{R}_T / |\underline{R}_T|$, etc., with c_0 = mean speed of sound, as shown in Fig. A.2-2. A second extension (FOM⁽²⁾) of the original FOM model, in addition to the above dopplerized delay (A.2-13), explicitly includes the tilted surface Σ of surface wave heights in the "cross-section" $\gamma_0^{(1)}$, e.g.,

$$\hat{\gamma}_{0}^{(1)} = \gamma_{0}^{(1)} \, \hat{\underline{n}}_{q} \cdot (\hat{\underline{i}}_{T} - \hat{\underline{i}}_{R}) = \gamma_{0}^{(1)} \, \hat{\underline{n}}_{q} \cdot (\hat{\underline{i}}_{0T} - \hat{\underline{i}}_{0R}), \tag{A.2-14}$$

where now $\hat{\underline{n}}_g$ is the normal to the (large-scale) component of the gravity-capillary wave surface

$$\hat{n}_{g} = \left\{ \frac{\hat{1}_{x}\zeta_{x} + \hat{1}_{y}\zeta_{y} - \hat{1}_{z}}{\sqrt{1 + \zeta_{x}^{2} + \zeta_{y}^{2}}} \right\}_{g} = \left\{ (\hat{1}_{x}\zeta_{x} + \hat{1}_{y}\zeta_{y} - \hat{1}_{z})n_{z} \right\}_{g}. \tag{A.2-15}$$

For both extensions $\Gamma^{(0)} = \sigma^{(1)}(\underline{r})\hat{\gamma}_0^{(1)}$, where again $\sigma^{(1)}$ is the mean density of (independent) scatterers on the surface. FOM^(1a) and FOM⁽²⁾ thus incorporate explicitly the physical geometry of the (wave) surface.

The practical utility of FOM models is that they replace the often very difficult problem of explicitly evaluating the boundary conditions at the scattering elements, by introducing a time-variable weighting function $\gamma_0^{(1)}h_F^{(1)}$, or $\hat{\gamma}_0^{(1)}h_F^{(1)}$, which also may include the extended doppler delay (A.2-13) [36]. Simple choices of $\gamma_0^{(1)}$, $h_F^{(1)}$ are then tested against experience [43]-[46]. The limitations of the FOM models are threefold:

- (1). They lack explicit physical structures. [This can be to a considerable extent overcome by using $FOM^{(2)}$ models, cf. above];
- (2). They are quasi-phenomenological, namely, they require a suitable choice of $h_F^{(1)}$ and consequent calibration to the problem at hand; and
- (3). They (currently) omit multiple-scatter effects, or diffraction (e.g., $v^{(k)} = 0$, $k \ge 2$). [This is also a limitation of most "classical" scatter models, as well [25]-[27], [37]-[39], [48], but see [51].]

B. Classical Scatter Models:

Classical scatter models cf. [25]-[27], [48], as noted in Sec. 5 earlier, in their practical formulations for the most part neglect multiple scatter effects ($k \ge 2$), so that, in our formalism, cf. Appendix A.1, $v^{(0)} = v^{(1)}$, $v^{(k)} = 0$, $k \ge 2$. In the important cases of surface scatter, for example, it can be shown on comparing the classical approximations, namely Kirchoff and perturbational boundary evaluations [48], that now

$$\gamma_{o} v^{(0)} \stackrel{:}{=} \gamma_{o} v^{(1)} \stackrel{:}{=} RS \frac{\omega}{c_{o}} \hat{n}_{g} \cdot (\hat{\underline{i}}_{T} - \hat{\underline{j}}_{R}), \qquad (A.2-16)$$

cf. (5.10) above, with (A.2-15). Equations (5.6)-(5.7c) again apply here, so that

$$K_{\hat{Q}|surface} = \mathbf{X} \{\langle [\hat{\underline{\mathbf{n}}}_{g} \cdot (\hat{\underline{\mathbf{i}}}_{T} - \hat{\underline{\mathbf{i}}}_{R})]_{1} [\hat{\underline{\mathbf{n}}}_{g} \cdot \hat{\underline{\mathbf{i}}}_{T} - \hat{\underline{\mathbf{i}}}_{R}]_{2} \rangle - \langle \hat{\mathbf{0}} \rangle_{1} \langle \hat{\mathbf{0}} \rangle_{2}$$
 (A.2-17)

for the covariance of the incoherent scatter from a typical random (wave) surface. A detailed evaluation of (A.2-16) is described in [50], Part I, and extended to include explicit doppler measures in Part II. Higher-order moments are found similarly, viz.

$$\langle \hat{Q}_1 \hat{Q}_2 \hat{Q}_3 \rangle_{\text{surface}} = \chi \{ \langle \gamma_{01} \gamma_{11}^{(1)} \gamma_{02} \gamma_{21}^{(1)} \gamma_{03} \gamma_{31}^{(1)} \rangle \}, \text{ etc.,}$$
 (A.2-18)

with the help of (A.2-16) explicitly, see [50].

Appendix A.3 Bayesian Estimation: A Decision-Theoretic Formulation

Here we briefly summarize the needed main elements of the decision-theoretic formulation of signal estimation theory [15], [58], [20].

For an estimator we write

$$\underline{Y} = \underline{Y}_{\Omega}(\underline{\theta} | \underline{X}) = g_{F}(\underline{X}), \qquad (A.3-1)$$

where $\underline{\theta}=(\theta_1,\ldots,\theta_{\widehat{M}})$ is a set of signal parameters (or waveform samples, etc.) to be estimated; $\underline{X}=(X_1,\ldots,X_J)$ is the set of received data samples in which the estimate is to be based; the "decisions," or estimates made are denoted by $\underline{Y}=(Y_1,\ldots,Y_{\widehat{M}})$, and σ in Y_{σ} indicates the apriori probability (density) governing the parameters $\underline{\theta}$, e.g., $\sigma=\sigma(\underline{\theta})$, or \underline{S} in $\sigma(\underline{S})$. [We note that \underline{Y}_{σ} is an estimator for all permitted \underline{X} , while for a particular set $\underline{X}=\underline{X}'$, Y_{σ} becomes an estimate.] The decision rule δ is here $\underline{\delta}(\underline{Y}|\underline{X})=\underline{\delta}(\underline{Y}-\underline{Y}_{\sigma}(\underline{\theta}|\underline{X}))$, which is an (M-dimensional) delta function. When \underline{Y}_{σ} (=g_E) is given, i.e., the estimating receiver $g_{E}(\underline{X})$ is specified, the average error (or risk) is determined from

$$R(\sigma,\delta)_{S \text{ or } \theta} = \int_{\Omega} \sigma(\underline{S} \text{ or } \underline{\theta}) \underline{dS} \text{ (or } \underline{d\theta}) \cdot \int_{\Gamma} F_{n}(\underline{X}|\underline{S}(\underline{\theta})) C(\underline{S} \text{ or } \underline{\theta},\underline{Y}_{\sigma}) \underline{dX}$$
(A.3-2)

where C is an appropriately chosen <u>cost function</u>, naturally proportional to the measure of error selected. The estimator \underline{Y}_{σ} is a <u>point estimate</u>, and \underline{F}_{n} is the conditional pdf of \underline{X} , given $\underline{S}(\underline{\theta})$.

Another useful type of estimation procedure is <u>interval estimation</u>, defined by the probability P that a particular point estimate, Y_{σ} (for given X) falls within (1 ± 1) 100% of the true value of the quantity $[\underline{S} = \text{waveform}, \text{ or } \underline{\theta} = \text{parameters in } \underline{S}(\underline{\theta})]$ being estimated, viz.:

$$P\{(1-\lambda)[\underline{S} \text{ or } \underline{\theta}] \leq Y_{\sigma}(\underline{S} \text{ or } \underline{\theta}|\underline{X}) \leq (1+\lambda)[\underline{S} \text{ or } \underline{\theta}]\}, \quad 0 < \lambda < 1$$
(A.3-3)

where λ is a measure of the prechosen <u>confidence interval</u> selected, $-\lambda \underline{S}$ to $+\lambda \underline{S}$ for example. In the case of a single parameter θ , for instance, we can write

$$P\{(1-\lambda)\theta \leq \Upsilon_{\sigma}(\theta | \underline{X}) \leq (\underline{1}+\lambda)\theta\} = \int_{(1-\lambda)\theta}^{(1+\lambda)\theta} W_{1}(\Upsilon_{\sigma}(\varepsilon))d\Upsilon_{\sigma}, \qquad (A.3-4)$$

where $W_1(\Upsilon_\sigma|\theta) = p(\Upsilon|\theta)$ is the conditional pdf of the estimator Υ , conditional on θ , formed here from $p(\Upsilon|\theta) = \int_T F_n(\underline{X}|\underline{S}(\theta)\delta(\Upsilon_\theta|\underline{X})\underline{dX})$, with obvious formal extensions to the multidimensional cases, cf. [16], [58], [20]. The unconditional pdfs of the estimators themselves are obtained from $p(\underline{\Upsilon}) = \langle p(\underline{\Upsilon}|\underline{S}) \text{ or } \underline{\theta} \rangle_{\underline{S}}$ or $\underline{\theta}$.

We emphasize that Υ_{σ} is a <u>point estimator</u>, embodying the specific structure, g_E , (A.3-1), of the receiver performing the estimation. On the other hand, interval estimators, as expressed by P, (A.3-3), (A.3-4), yield a probability which is a measure of the efficiency of the point estimator for any particular application (i.e., choice of χ). The average error (or risk) (A.3-2) measures the expected cost or average error in using Υ_{σ} , considered over all possible $\{\chi\}$ received.

For optimum, or Bayes estimation we seek estimators $\underline{\gamma}^\star_\sigma$ which minimize the average error or risk R(σ , δ), (A.3-2). The general form of the resulting $\underline{\gamma}^\star_\sigma$ depends, of course, on the choice of "cost" or error function C(\underline{S} or $\underline{\theta},\underline{\gamma}_\sigma$). For example, for the quadratic cost function (QCF)

$$C(\underline{\theta},\underline{\gamma}_{\sigma}) = C_{0} |\underline{\theta}-\underline{\gamma}_{\sigma}|^{2} = C_{0} |\underline{\hat{\theta}}-\underline{\gamma}_{\sigma}|^{2}, \qquad (A.3-5)$$

the associated optimum estimator is found to be (cf. Chapter 3, [58], Chapter 21, [20]) the set of equations

$$\frac{\chi_{\sigma}^{*}(\underline{\theta}|\underline{X})_{QCF} = \int_{\Omega_{\underline{\theta}}} \underline{\underline{\vartheta}}\sigma(\underline{\underline{\vartheta}})F_{n}(\underline{X}|\underline{\underline{S}}(\underline{\theta}))\underline{d}\underline{\theta} / \int_{\Omega_{\underline{\theta}}} \sigma(\underline{\underline{\theta}})F_{n}(\underline{X}|\underline{\underline{S}}(\underline{\underline{\vartheta}}))\underline{d}\underline{\theta}, \qquad (A.3-6)$$

with $\underline{\sigma} + \underline{S}$, $\sigma(\underline{\theta}) + \sigma(\underline{S})$, $\underline{d\theta} + \underline{dS}$ in the case of estimating signal waveforms. Note that $\underline{\Upsilon}_{\sigma}^*$ is generally a <u>nonlinear</u> operator on the received data, \underline{X} .

Another cost function of considerable interest is the <u>simple</u> (or rectangular) cost function (SCF) given by

$$C(\underline{\hat{\varphi}},\underline{\hat{\gamma}}_{\sigma}) = C_{o} \sum_{\hat{m}=1}^{\hat{M}} [A_{\hat{m}} - \hat{\gamma}(\hat{\gamma}_{\hat{m}} - \theta_{\hat{m}})]$$
(A.3-7)

with appropriate choices of C_0 , $A_{\widehat{m}}$ to ensure meaningful results (e.g., positive errors, etc.). Minimization of average risk here leads directly to (cf. Sec. 21.2-1 of [20]) the following relations determining $Y_{\sigma}^{\star}|_{SCF}$:

$$\frac{\partial}{\partial \theta_{\widehat{\mathbf{m}}}} \log \left\{ \sigma(\theta_{\widehat{\mathbf{m}}}) \mathsf{W}_{\mathbf{J}}(\underline{\mathsf{X}}|\theta_{\widehat{\mathbf{m}}}) \right\} \Big|_{\theta_{\widehat{\mathbf{m}}} = \theta_{\widehat{\mathbf{m}}}^{*} = \gamma_{\widehat{\mathbf{m}}}^{*}} = 0, \quad \text{all } \widehat{\mathbf{m}} = 1, \dots, \widehat{\mathsf{M}},$$
 (A.3-8)

where $W_{J}(\underline{X}|\gamma_{\widehat{m}}=\theta_{\widehat{m}}) = \langle F_{n}(\underline{X}|\underline{S}(\gamma_{\widehat{m}},\underline{\theta}')) \rangle_{\underline{\theta}'=\underline{\theta}-\underline{\gamma}_{\widehat{m}}} / \sigma(\gamma_{\widehat{m}}=\theta_{\widehat{m}})$, and

$$\begin{split} &\sigma(\gamma_{\widehat{m}}=\theta_{\widehat{m}})=<\sigma(\gamma_{\widehat{m}},\underline{\theta'})>_{\underline{\theta'}}, \text{ with }\underline{\theta'}=\text{all }\underline{\theta}\text{ except }\underline{\theta}_{\widehat{m}}. \text{ The condition (A.3-8)}\\ &\text{determining }\underline{\gamma}^*=\left\{\gamma_{\widehat{m}}^*\right\} \text{ is precisely that determining the }\underbrace{\text{unconditional}}\\ &\underset{maximum \ likelihood \ estimates}{\text{estimates}} \text{ (UMLE's) of }\theta=\left\{\theta_{\widehat{m}}\right\}, \text{ viz. }\gamma_{\widehat{m}}^*\text{ here. } \text{ [If }\sigma(\theta_{\widehat{m}})\text{ is omitted in (A.3-8) one has the corresponding }\underbrace{\text{conditional }maximum\\ \text{likelihood estimates (CMLE's) of the }\left\{\theta_{\widehat{m}}^*\right\}. \text{ The extension of (A.3-8) to}\\ &\text{waveform estimation is formally immediate: one replaces }\theta_{\widehat{m}}\text{ by }S_{\widehat{m}}, \sigma(\theta_{\widehat{m}})\\ &\text{by }\sigma(S_{\widehat{m}}), \text{ $W_{J}(\underline{X}|\theta_{\widehat{m}})$ by $W_{J}(\underline{X}|S_{\widehat{m}})$, etc., with $\widehat{m}=1,\ldots,J$ now.]} \end{split}$$

In summary, we remark that:

- I. The maximum (conditional) likelihood estimator (CMLE) maximizes the probability of a correct decision, without regard to incorrect decisions and their costs;
- II. The maximum unconditional likelihood estimator (UMLE) maximizes the probability of a correct estimate, when all possible (signal or parameter) values are taken into account, again without specific regard for incorrect decisions and their errors;

>(A.3-9)

III. The optimum quadratic estimator Υ_{0CF}^{\star} , (A.3-6), is an unconditional least mean-square error (LMSE) estimator, which accounts on the average for incorrect decisions and errors.

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